

Experimental Tests of CPT Invariance at CERN

Szegedi Egyetem, Elm-Fiz. Tanszék, 2018.10.15.

Horváth Dezső

horvath.dezso@wigner.mta.hu

Wigner Research Centre for Physics,
Institute for Particle and Nuclear Physics, Budapest, Hungary

&

ATOMKI, Debrecen, Hungary



Outline

- Antimatter and its lack in the Universe
- CPT invariance: matter–antimatter symmetry
- The Antiproton Decelerator at CERN
- Antimatter experiments at CERN
- Antihydrogen: production
- Antihydrogen: spectroscopy
- Antihydrogen: gravity
- Outlook: ELENA
- Use of antimatter in life

Birth of antimatter

Paul Dirac, 1928: Linear equation for the hydrogen atom.
Square root of a quadratic equation \Rightarrow two solutions for electrons ($x^2 = 4 \Rightarrow x = \pm 2$).

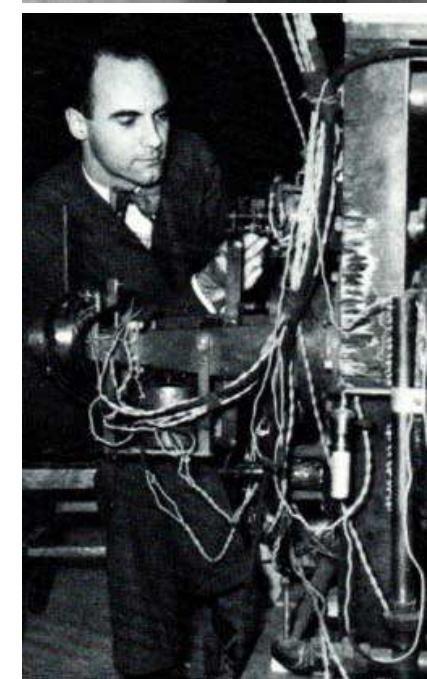
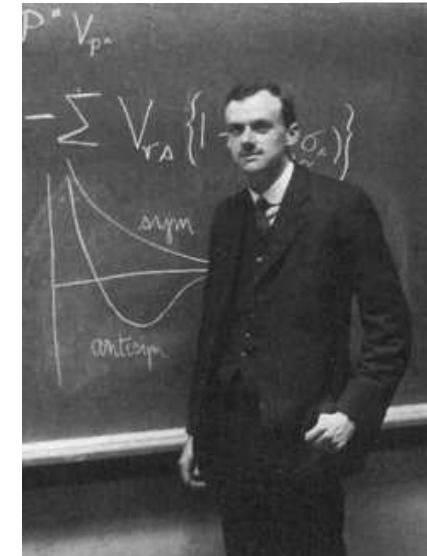
Two kinds of electrons:

- + mass and — charge (ordinary electron);
- — mass and + charge (anti-electron = positron).

Negative mass non-physical. Dirac: particle holes.

Carl Anderson (1932): e^+ in cosmic rays!
 \Rightarrow real existing particle: positron.

Nobel prizes (in 4 years): Dirac: 1933; Anderson: 1936



Antiparticles

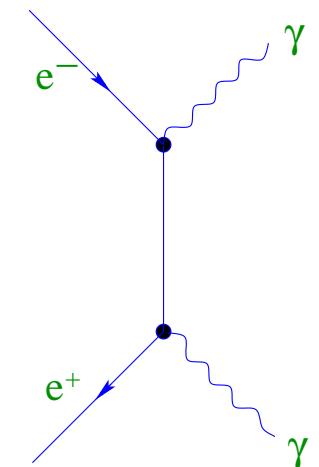
Every matter particle (fermion) has an antiparticle

Proton (hydrogen nucleus) \leftrightarrow antiproton.

Particles and antiparticles must have the same properties
apart from the signs of charges.

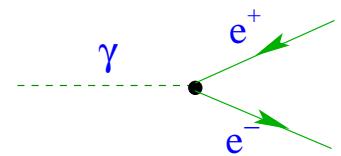
When particle meets its own antiparticle they annihilate to
photons or to lighter particles (energy conservation).

A slow positron in matter annihilates with an atomic electron by
emitting two (or three) gamma photons.



Converse reaction: radiation in the field of atomic nucleus can
produce particle + antiparticle pairs.

Low energy: $e^- + e^+$, higher energy ($E > 2M_p$): $p + \bar{p}$.



Antimatter mysteries

- Why there is practically no antimatter in our Universe?
At the Big Bang particles and antiparticles should have been produced together. Where did antimatter go?
- Could they be hiding in parts of the Universe inaccessible for us?
- Could there be a tiny difference between particle and antiparticle to cause this asymmetry?
- Are there particles which are their own antiparticles (Majorana particles)? Could the dark matter of the Universe consist of such particles?
- Can antimatter be used for something in everyday life or is it just an expensive curiosity?

Matter–antimatter symmetry

Charge conjugation: $C|\mathbf{p}(r, t)\rangle = |\bar{\mathbf{p}}(r, t)\rangle$

CPT invariance Space reflection: $P|\mathbf{p}(r, t)\rangle = |\mathbf{p}(-r, t)\rangle$

Time reversal: $T|\mathbf{p}(r, t)\rangle = |\mathbf{p}(r, -t)\rangle K$

(K : complex conjugation for $\exp\{-iEt\}$)

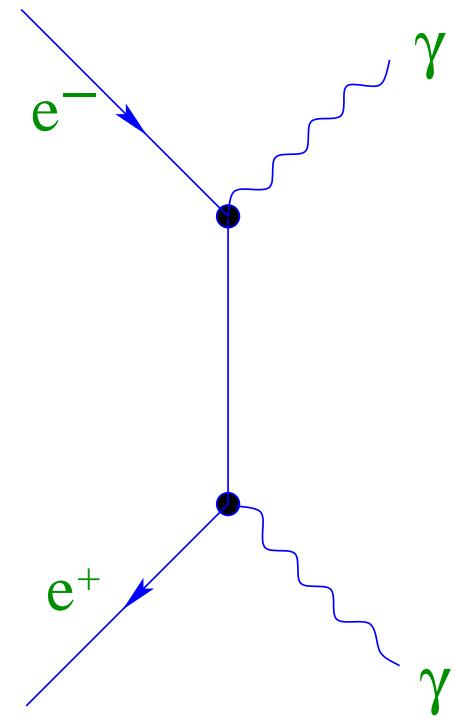
Basic assumption of field theory:

$$CPT|\mathbf{p}(r, t)\rangle = |\bar{\mathbf{p}}(-r, -t)\rangle \sim |\mathbf{p}(r, t)\rangle$$

meaning free antiparticle \sim particle
going backwards in space and time.

Giving up *CPT* one has to give up:

- locality of interactions \Rightarrow causality, or
- unitarity \Rightarrow conservation of matter, information, ...
- or Lorentz invariance



CPT-violating theories

Weak interaction violates P and CP symmetry

Theoreticians in general: CPT cannot be violated

- Standard Model is valid up to Planck scale ($\sim 10^{19}$ GeV).
Above Planck scale new physics \Rightarrow Lorentz violation possible
- Quantum gravity: fluctuations \Rightarrow Lorentz violation
Loss of information in black holes \Rightarrow unitarity violation

Motivation for testing CPT at low energy

- Quantitative expression of Lorentz and CPT invariance needs violating theory
- low-energy tests can limit possible high energy violation

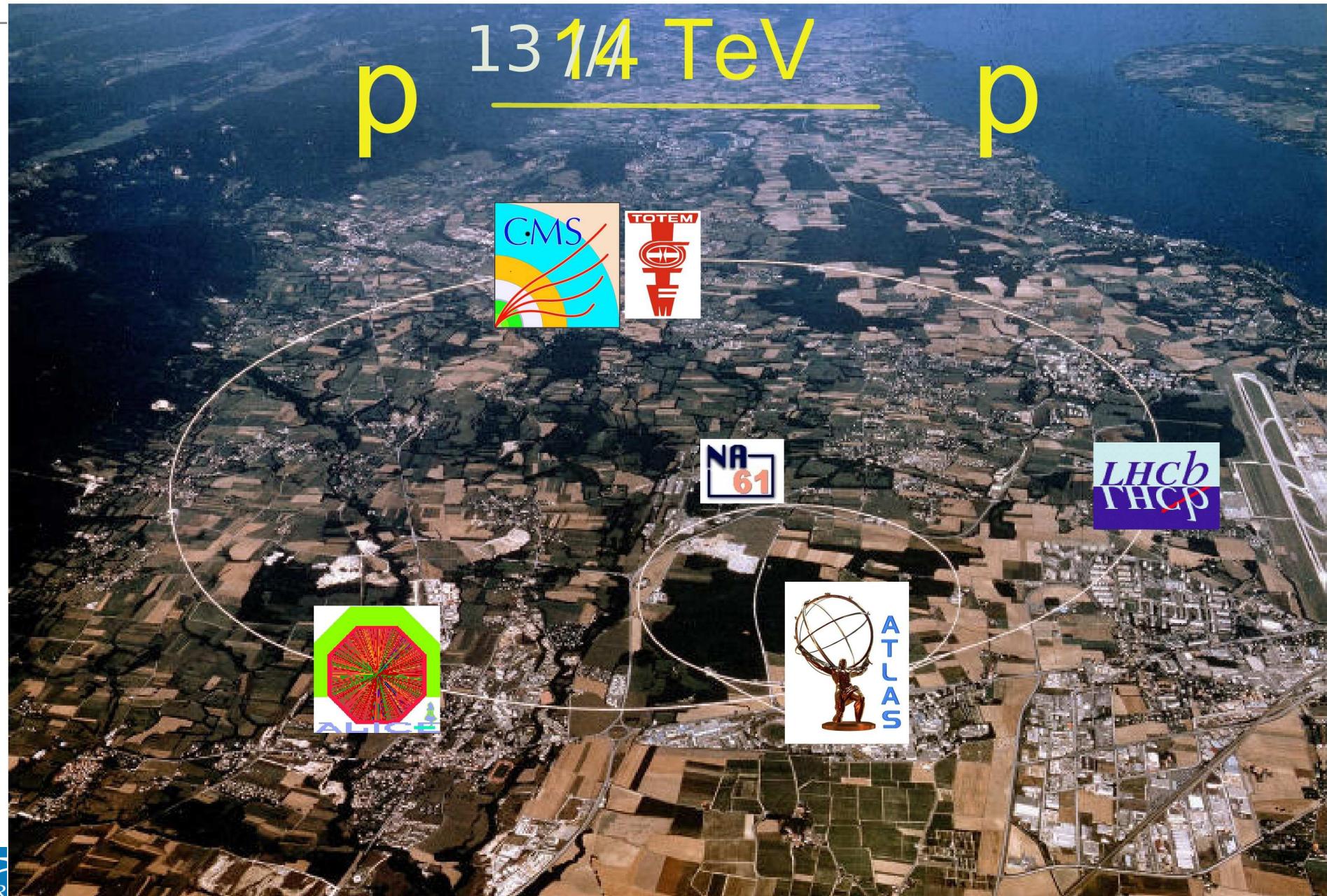
(Alan Kostelecký, F.R. Klinkhamer, N.E. Mavromatos et al)

How to test CPT ?

Particle = – antiparticle ?

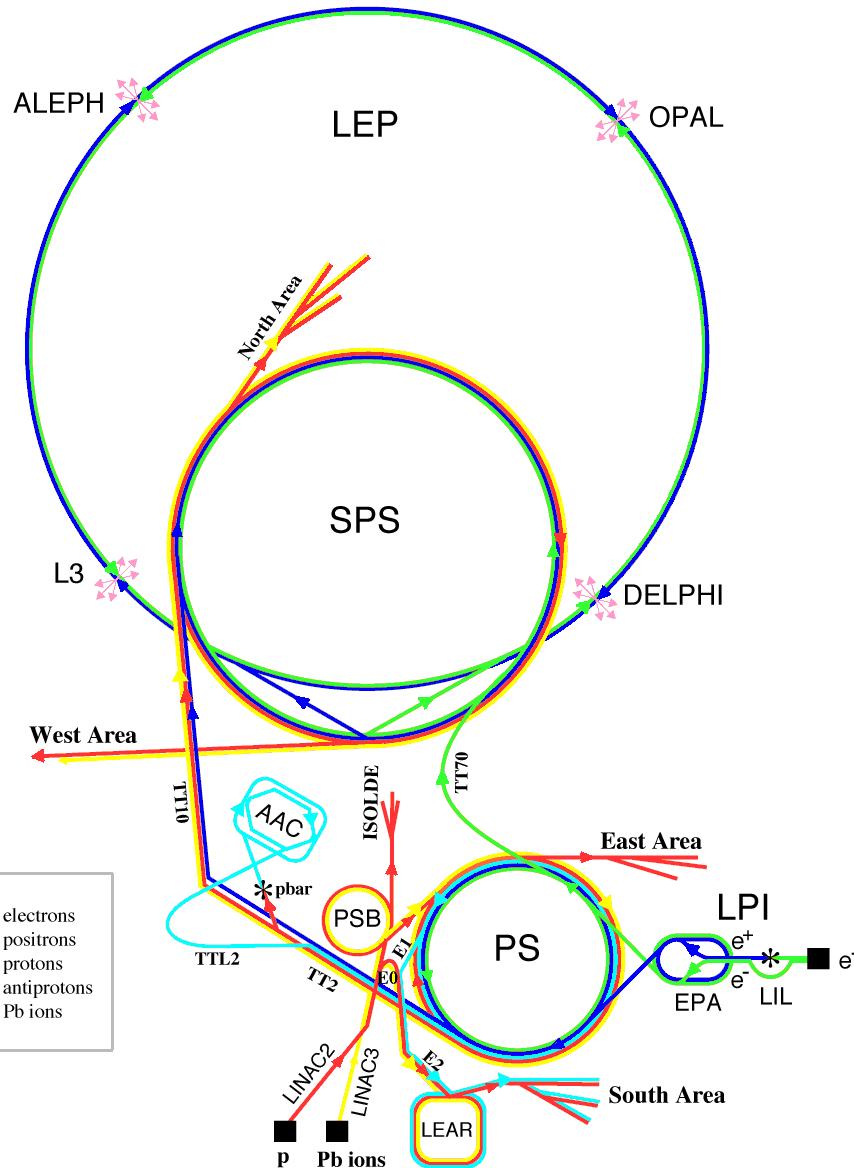
- $[m(K^0) - m(\bar{K}^0)]/m(\text{average}) < 10^{-18}$
- proton ~ antiproton? (compare $m, q, \vec{\mu}$)
- hydrogen ~ antihydrogen ($\bar{p}e^+$)? $2S - 1S$, HFS

CERN: aerial view

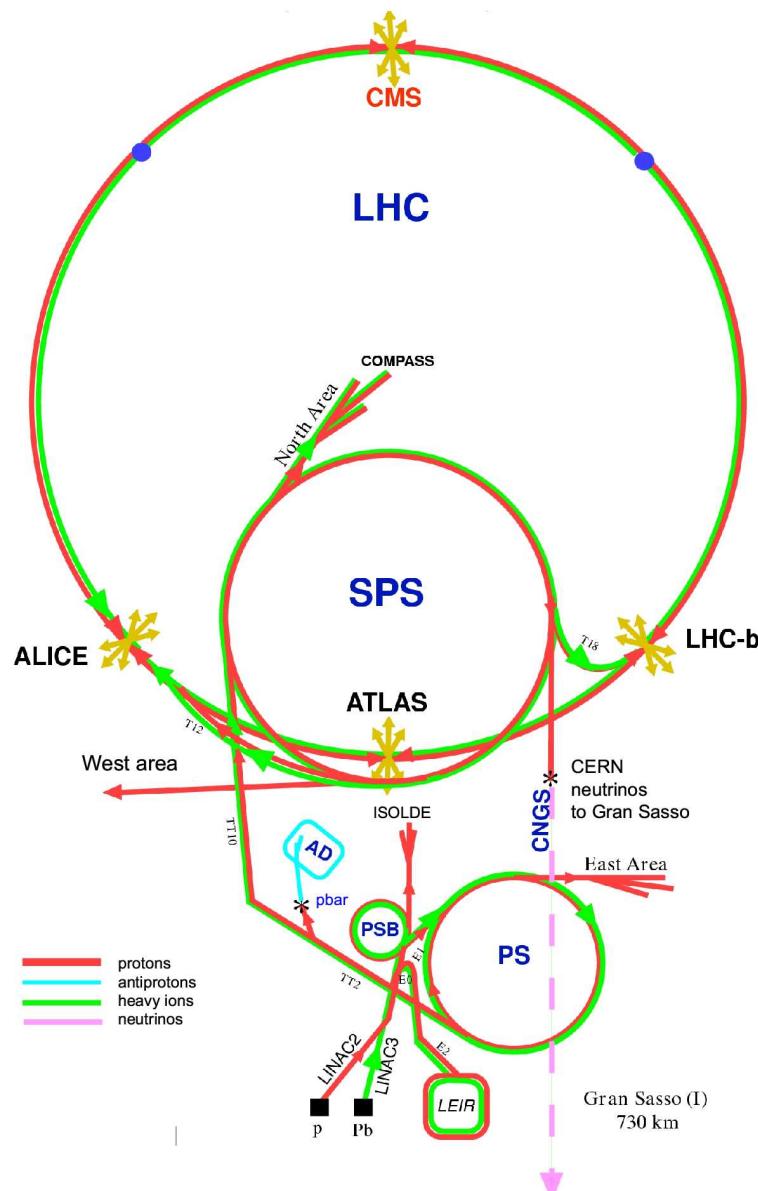


Accelerators at CERN

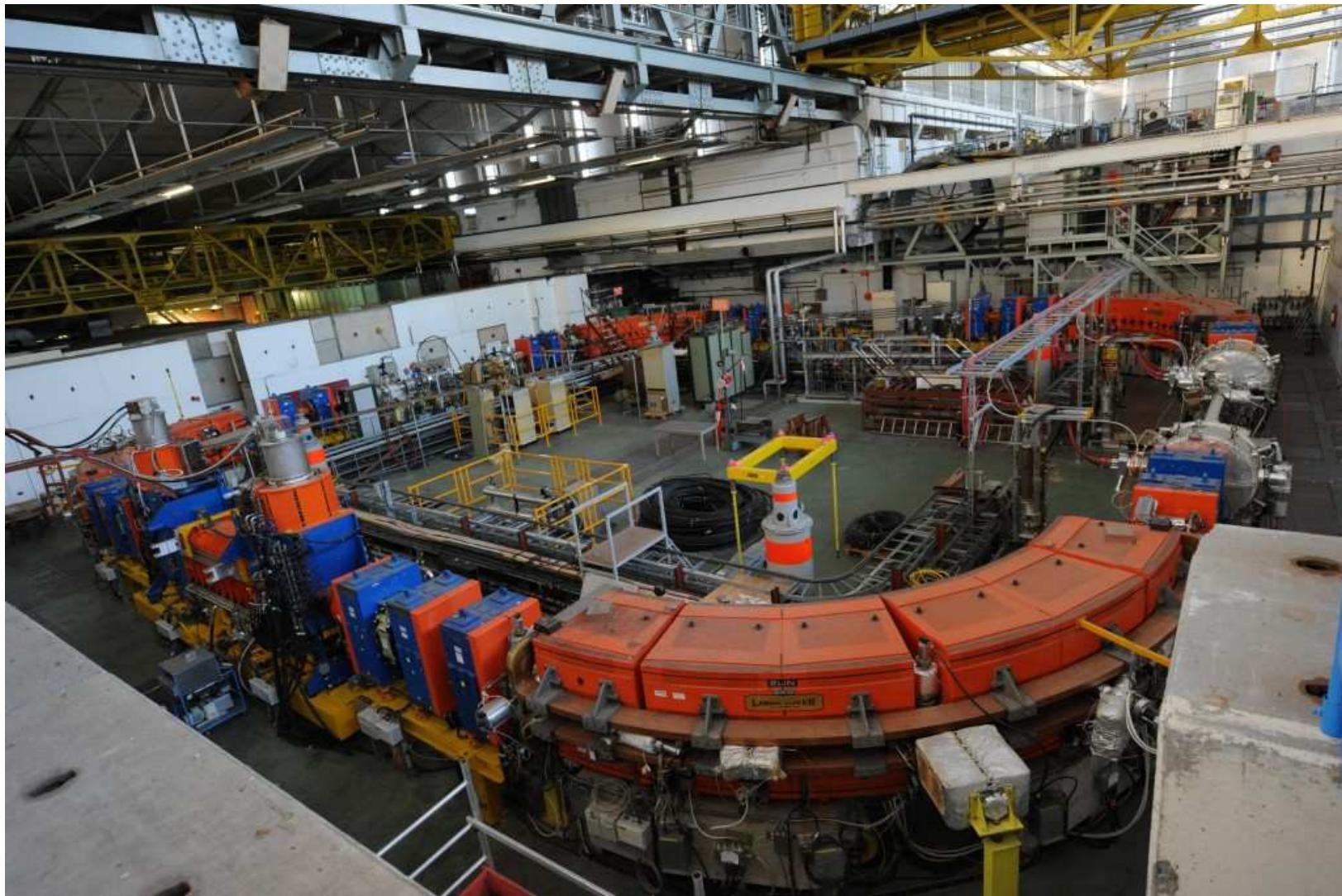
1989–2000



2009–2025??



Low Energy Antiproton Ring (LEAR)



1992: CERN wants to stop it
when **CLEAR** finishes data taking.



Save LEAR! Munich workshop, 1992

1993

ANTIHYDROGEN

T.W. Hänsch (Munich)

International Advisory Committee

W. Paul (Bonn)
J. Eades (CERN)
H.J. Kluge (Mainz)
R. Landua (CERN)
B.I. Deutch (Aarhus)
A.P. Mills (Bell Labs)
D. Möhl (CERN)
G. Gabrielse (Harvard)
M.H. Holzscheiter (Los Alamos)
R.J. Hughes (Los Alamos)
H. Kalinowsky (Mainz)
D. Kleppner (MIT)

H.J. Kluge (Mainz)
R. Landua (CERN)
A.P. Mills (Bell Labs)
D. Möhl (CERN)
W.D. Phillips (NBS)
G. Torelli (Pisa)
J.T.M. Walraven (Amsterdam)
T. Yamazaki (Tokyo)

Proceedings of the
Antihydrogen Workshop

Ludwig Maximilian University
Munich, Germany
30-31 July 1992



J.C. Baltzer AG

SCIENCE PUBLISHERS
BASEL - SWITZERLAND

Editor:
J. Eades
CERN
Geneva, Switzerland

Workshop Chairman

T.W. Hänsch (Munich)

International Advisory Committee

W. Paul (Bonn)
J. Eades (CERN)
H.J. Kluge (Mainz)
R. Landua (CERN)
B.I. Deutch (Aarhus)
A.P. Mills (Bell Labs)
D. Möhl (CERN)
G. Gabrielse (Harvard)
M.H. Holzscheiter (Los Alamos)
R.J. Hughes (Los Alamos)
H. Kalinowsky (Mainz)
D. Kleppner (MIT)

H.J. Kluge (Mainz)
R. Landua (CERN)
A.P. Mills (Bell Labs)
D. Möhl (CERN)
W.D. Phillips (NBS)
G. Torelli (Pisa)
J.T.M. Walraven (Amsterdam)
T. Yamazaki (Tokyo)

Program Subcommittee

J. Eades (CERN)
G. Gabrielse (Harvard)
T.W. Hänsch (Munich)
H.J. Kluge (Mainz)
D. Möhl (CERN)
J.T.M. Walraven (Amsterdam)

Organising Committee

J. Eades (CERN)
R. Landua (CERN)
C. Zimmermann (Munich)

Chairmen

W. Paul (Why antihydrogen?)
D. Möhl (Ultra low energy antiprotons)
J. Eades (Positron accumulation and positronium)
G. Torelli (Routes to antihydrogen)
C. Zimmermann (Trapping and spectroscopy of antihydrogen)



Save LEAR! feasibility study, 1992

EUROPEAN ORGANIZATION FOR NUCLEAR RESEARCH



CERN LIBRARIES, GENEVA



CM-P00043862

CERN SPSLC 92-45
SPSLC M-505
August 27, 1992

MEMORANDUM TO THE CERN SPSL COMMITTEE

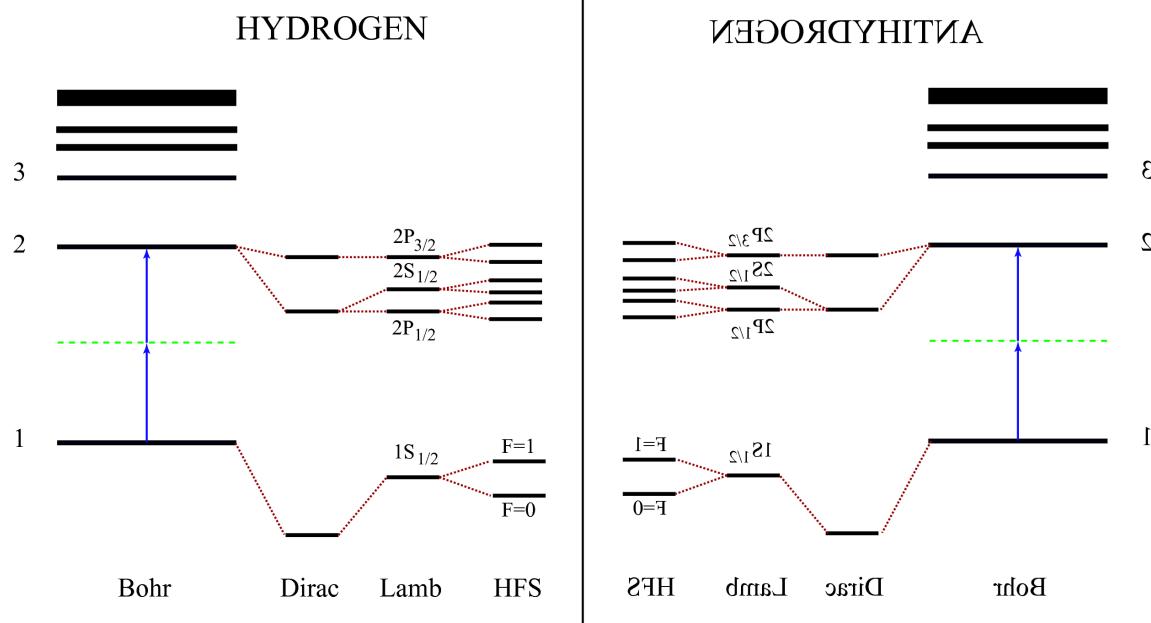
ANTIHYDROGEN FORMATION AND SPECTROSCOPY AT LEAR

M. Charlton¹⁾, J. Eades²⁾, D. Horváth^{3,4)} and R. J. Hughes⁵⁾

CERN extended LEAR until 1996
(when CPLEAR finished data taking).



Antihydrogen, $e^+ - \bar{p}$ atom, 1993



$2S - 1S$ transition
with 2-photons

Long lifetime,
narrow transition,
Doppler-free
spectroscopy

Feasibility study for the SPSL Committee of CERN (1992) converted into

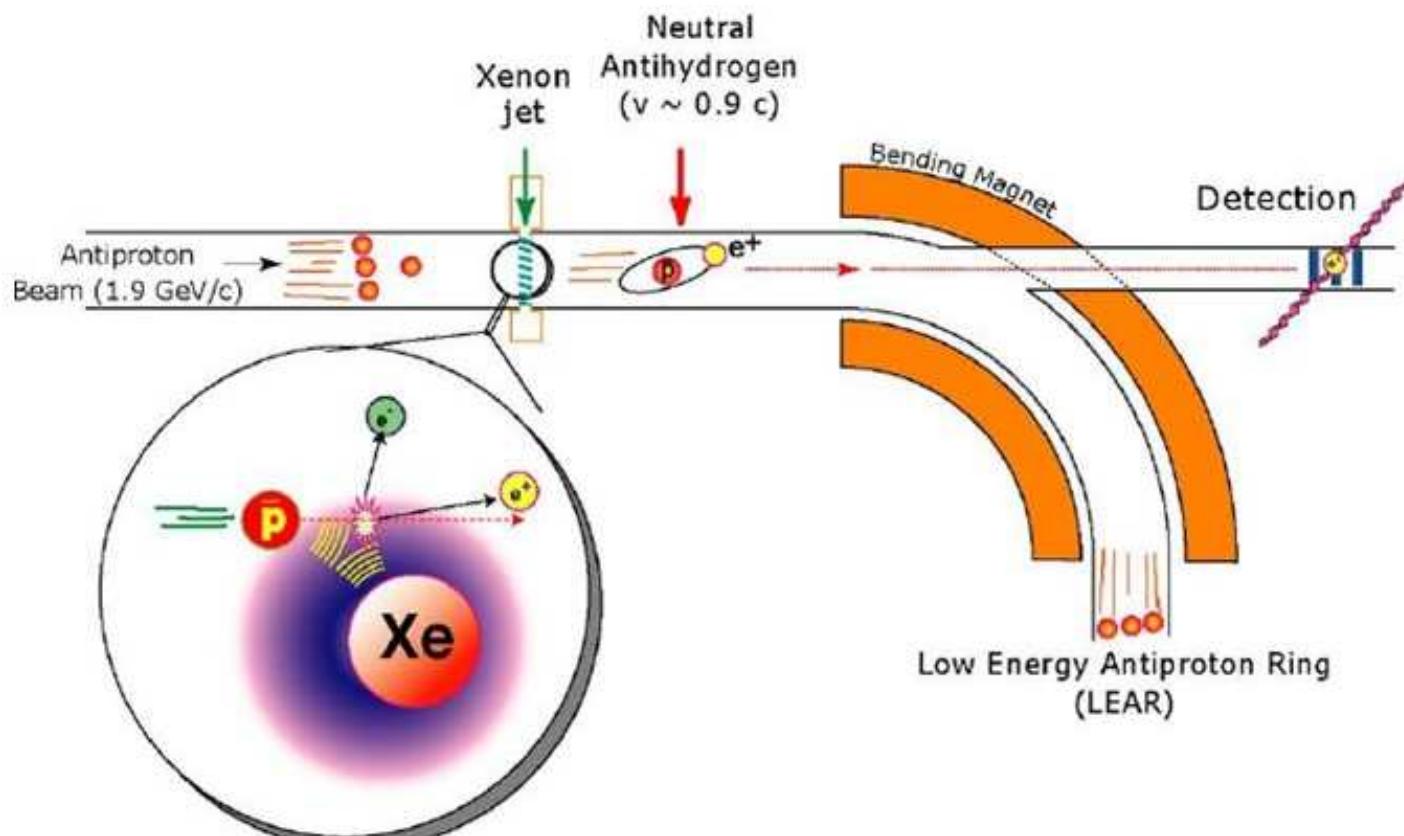
M. Charlton, J. Eades, D. Horváth, R. J. Hughes, C. Zimmermann:
Antihydrogen physics, Physics Reports 241 (1994) 65.

SPSLC accepted and CERN approved to build the Antiproton Decelerator

Great technical accomplishment of Dieter Möhl et al.



First (9) relativistic \bar{H} atoms at LEAR



G. Baur *et al.*, „Production of anti–hydrogen,” Phys. Lett. B 368 (1996) 251.

Later also at FERMILAB:

G. Blanford *et al.*, „Observation of atomic anti-hydrogen,”
Phys. Rev. Lett. 80 (1998) 3037.



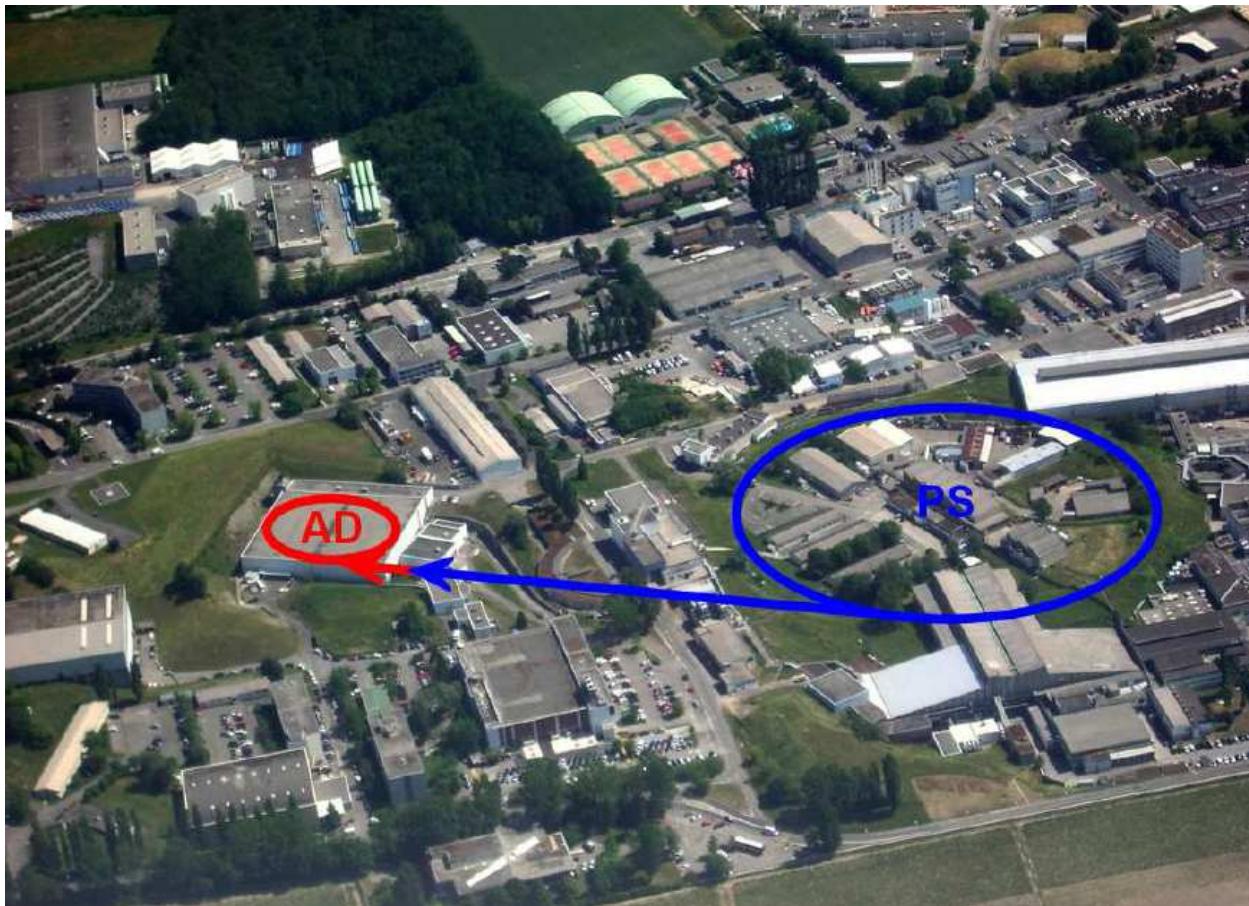
Antimatter factory at CERN



The Antiproton Decelerator at CERN

was built in 1997-99 to study antimatter physics

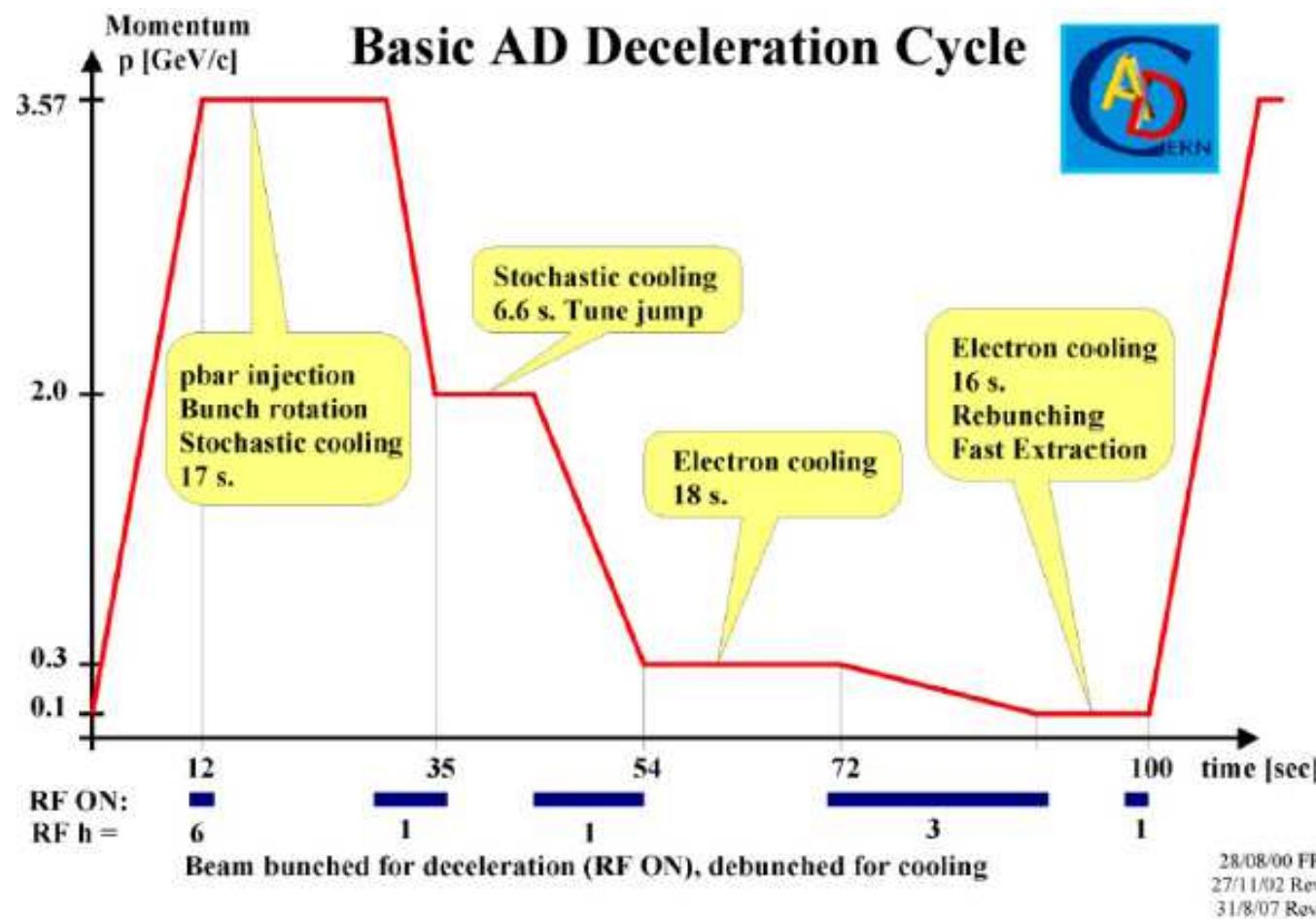
6 expts (3 each) for *CPT* and antigravity



©Ryugo S. Hayano, Tokyo U.



The Antiproton Decelerator: cooling



$\sim 4 \times 10^7$ 100 MeV/c antiprotons every 85 s

Pavel Belochitskii: AIP Conf. Proc. 821 (2006) 48



Steps toward \bar{H} spectroscopy

- Putting antiprotons (\bar{p}) in electromagnetic trap
- Trapping and cooling antiprotons
- Cooling slow positrons (e^+ from ^{22}Na) in trap
- Mixing \bar{p} and e^+ → recombination in $e^+e^+\bar{p}$ collisions (G. Gabrielse, ATRAP & Harvard U.)
- Trapping antihydrogen, waiting for deexcitation
- Cooling antihydrogen
- Laser spectroscopy on antihydrogen

2017: done by the ALPHA Collaboration!

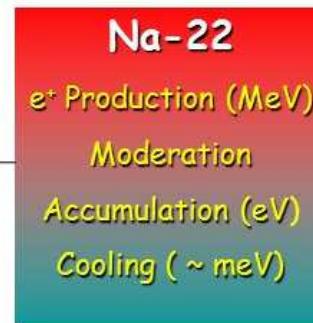
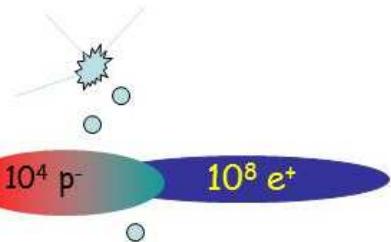
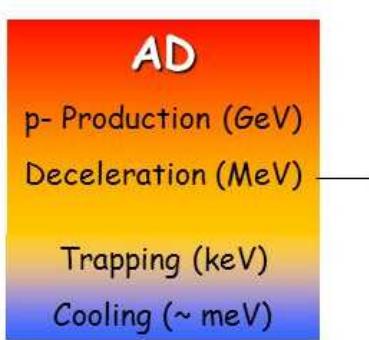


Antiproton production



CERN exhibition in Globe: \bar{p} production target at AD

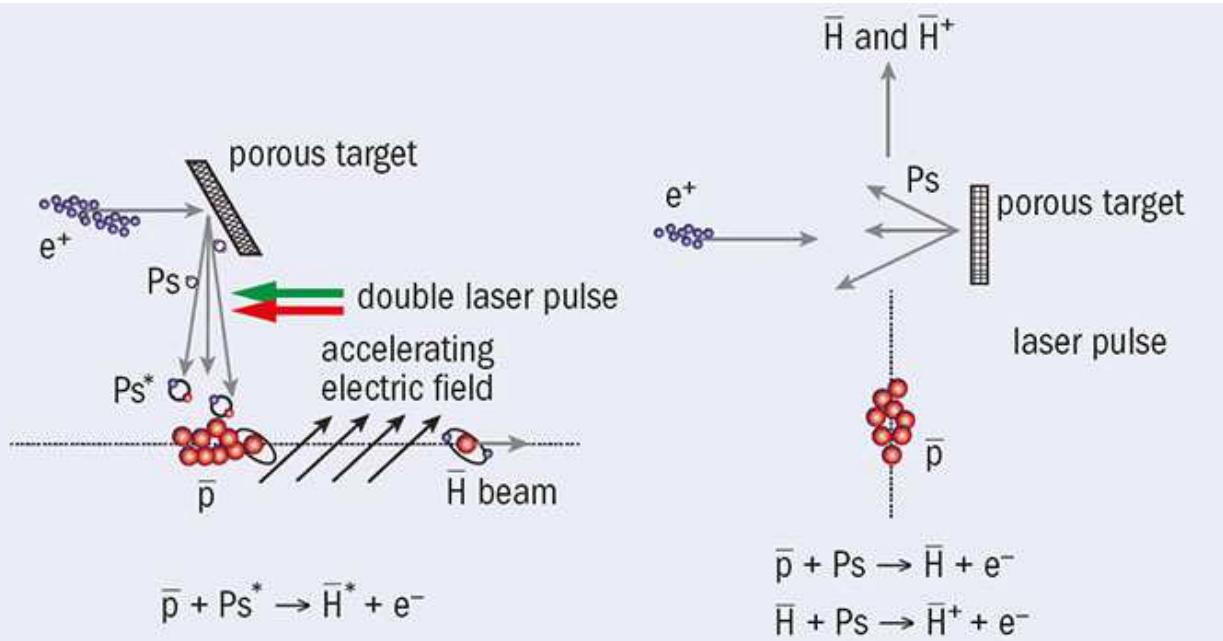
How to produce antihydrogen?



Radiative ($\bar{p}e^+\gamma$): deep bonding, low rate
(hopeless)

3-body ($\bar{p}e^+e^+$): shallow bond, high rate

Proposed by G. Gabrielse,
ATRAP & Harvard U.



With excited positronium:
high rate, deep bond
(planned)

Proposed by B. Deutch et al., Aarhus

ATHENA: first cold $\bar{\text{H}}$ atoms at AD

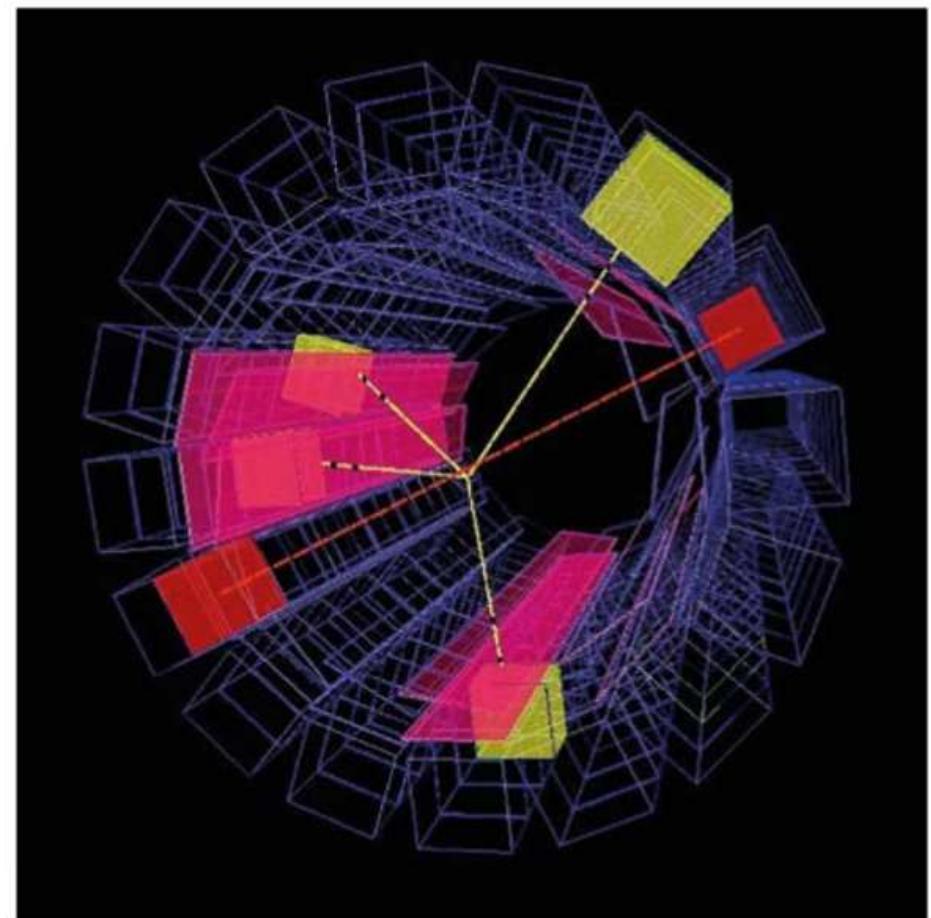
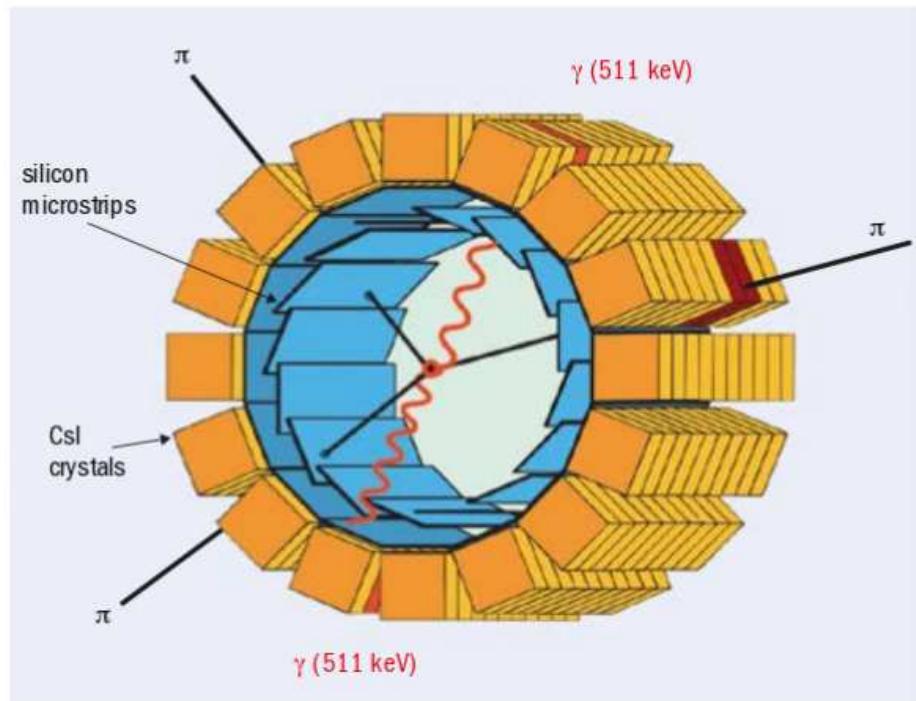


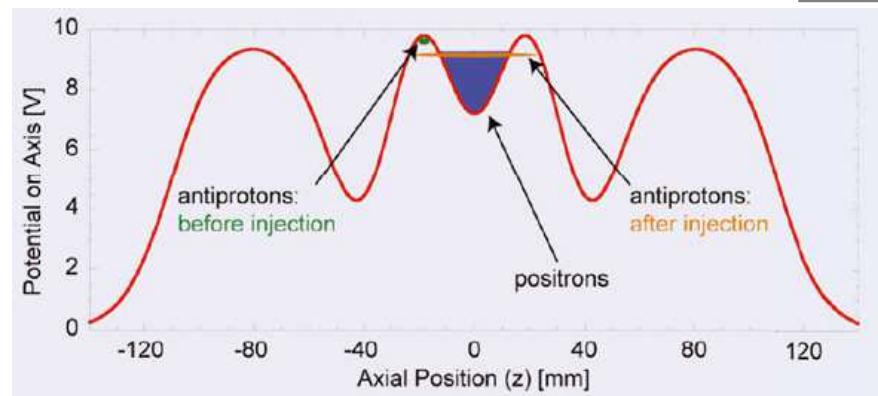
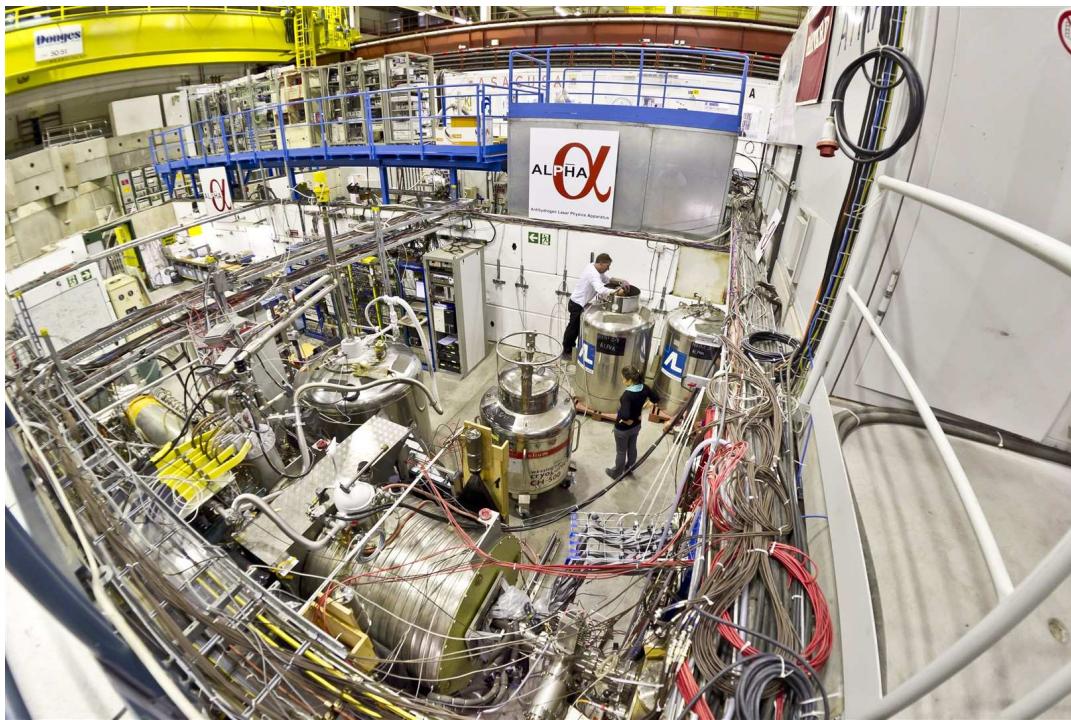
Fig.5. Above: A diagram of the ATHENA antihydrogen detector.
Right: An antihydrogen annihilation event in ATHENA,
reconstructing four charged pions (yellow) and two 511 keV
photons (red). (Image credits: ATHENA Collaboration.)

ATHENA Collaboration (1997 – 2005) \Rightarrow ALPHA Collaboration



ALPHA: \bar{H} production

ALPHA: Antimatter Laser PHysics
Apparatus (19 institutes of 9 countries)



- Capture 90,000 antiprotons.
- Mix with 3 million positrons.
- Produce 50,000 \bar{H} atoms.
- Remove charged particles.
- Trap 20 \bar{H} at $T = 0.54$ K.

\bar{H} kept trapped for 10 s \Rightarrow waiting deexcitation to $1S$ ground state.

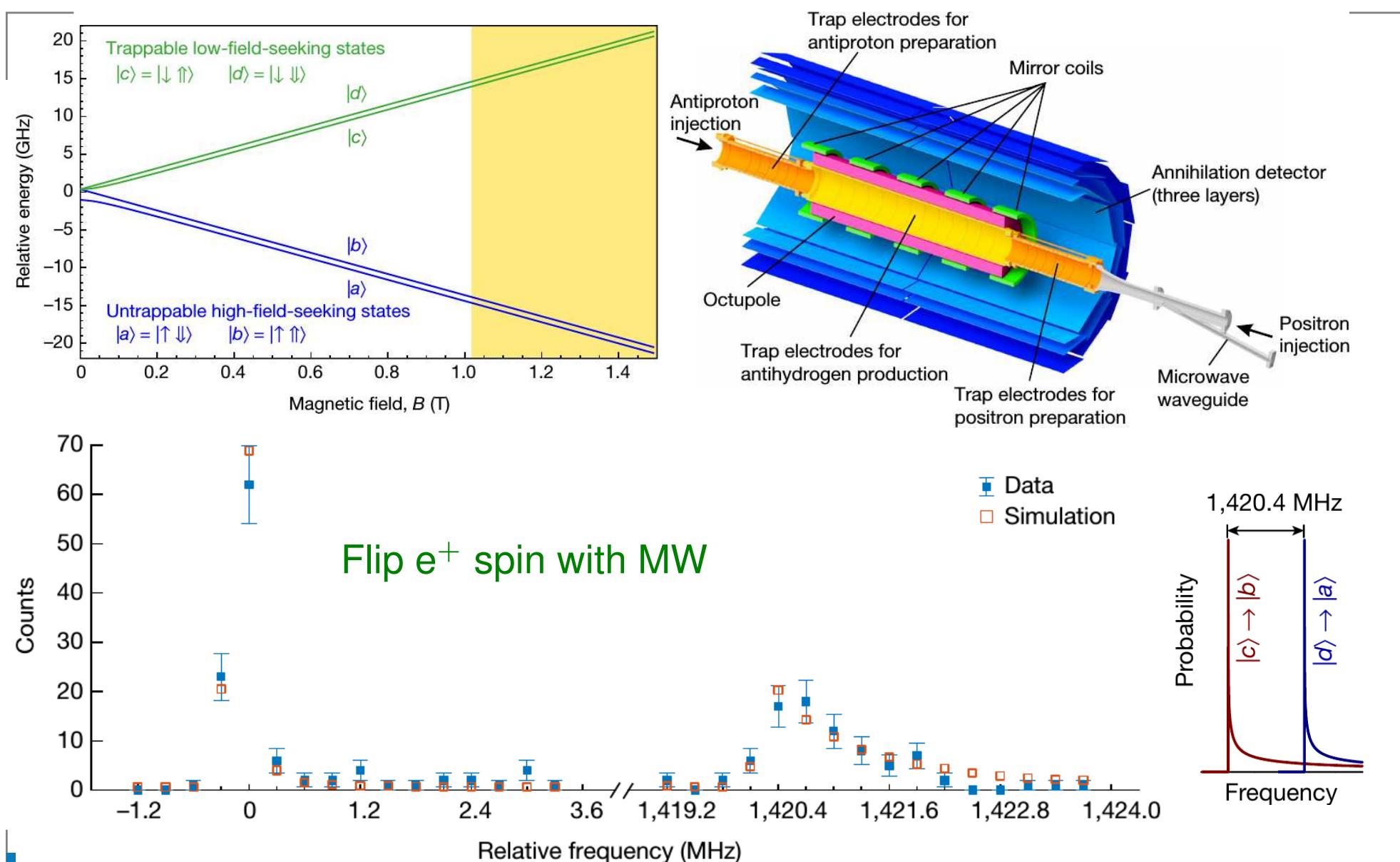
Demonstrated by keeping \bar{H} for >60 hours.

Detected and measured by dropping $B = 1$ T \Rightarrow annihilation.

Measured magnetic moment by hyperfine splitting on \bar{H}

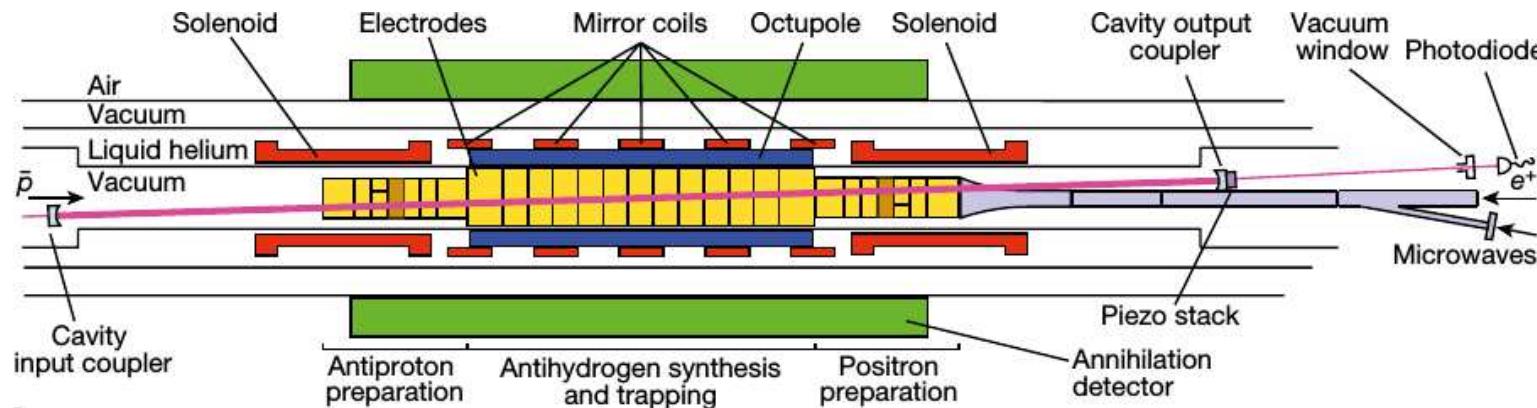


ALPHA: \bar{H} hyperfine spectrum



ALPHA Coll., „Observation of the hyperfine spectrum of antihydrogen,” Nature 548 (2017) 66.

ALPHA: \bar{H} $1S - 2S$ transition



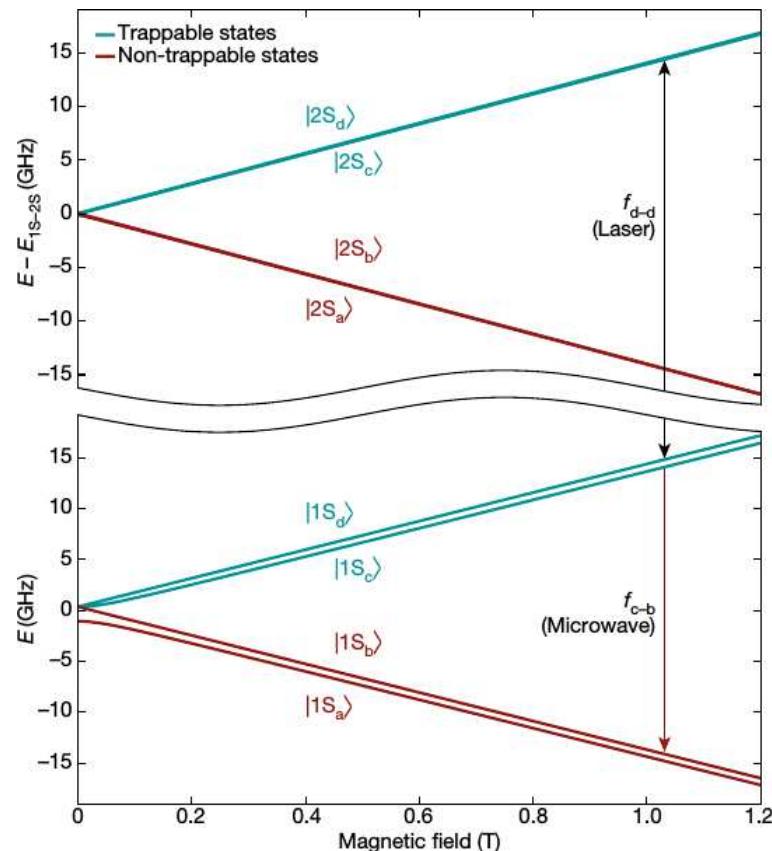
Measure annihilation rates.

Wait for 10 s to reach $\bar{H}(1S)$ state.

Excite $1S \rightarrow 2S$ with two 243 nm photons
(standing wave for 300 s) tuned around
resonance (appearance).

Use microwave to remove residual $1S$ atoms
(disappearance).

Flush trap by dropping B (residuals).



\bar{H} $1S - 2S$ spectroscopy

Result using 15000 \bar{H} atoms:

$$f_{d-d} = 2,466,061,103,079.4 \pm 5.4 \text{ kHz}$$

For hydrogen:

$$f_{d-d} = 2,466,061,103,080.3 \pm 0.6 \text{ kHz}$$

Difference (*CPT* test): 2×10^{-12}

ALPHA Coll.,

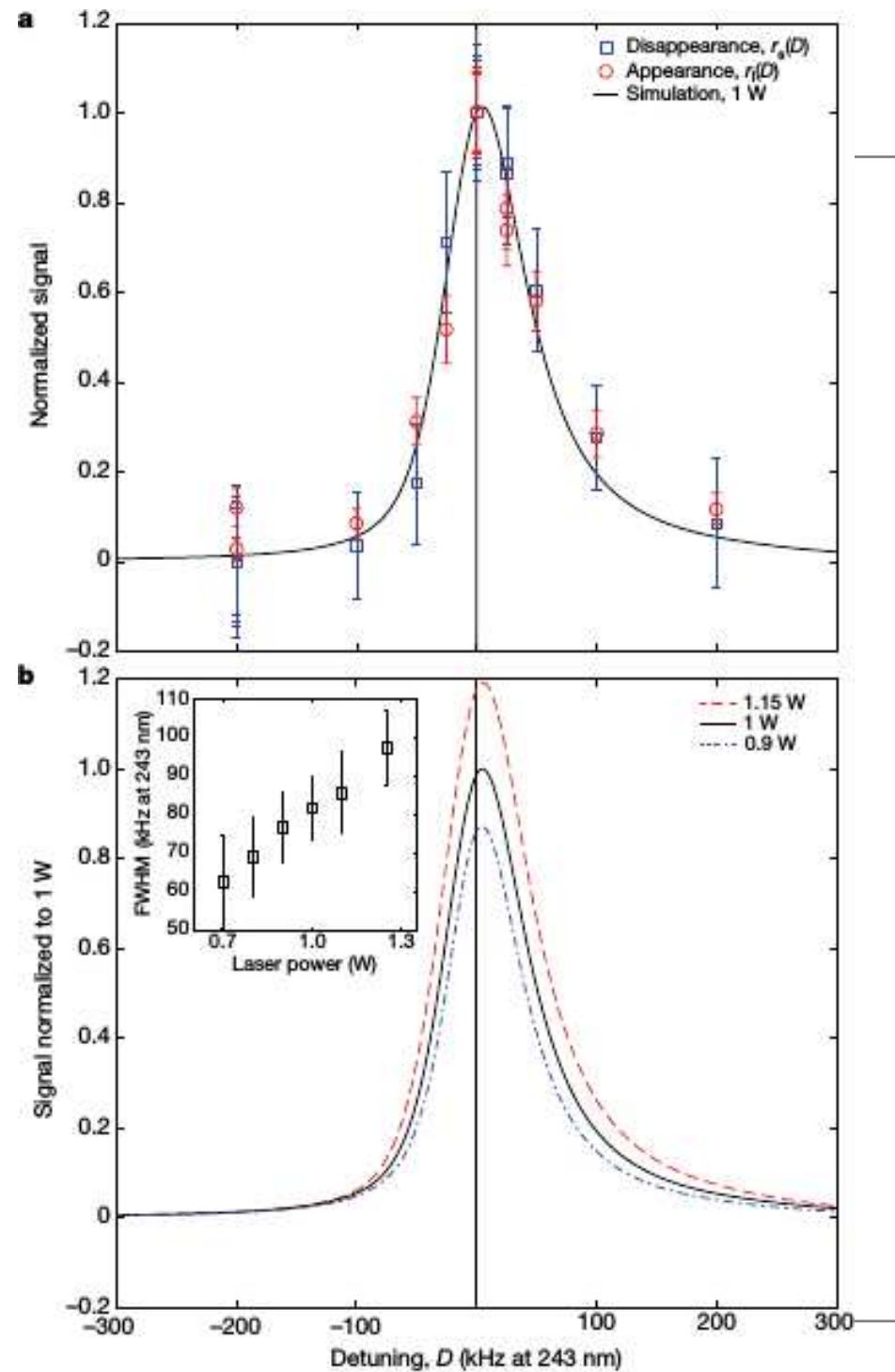
Observation of the $1S-2S$ transition in trapped antihydrogen,

Nature 541 (2017) 506.

ALPHA Coll.,

Characterization of the $1S-2S$ transition in antihydrogen,

Nature 557 (2018) 74.



ALPHA: \bar{H} $1S - 2S$ transition

ALPHA Collaboration (49 authors),

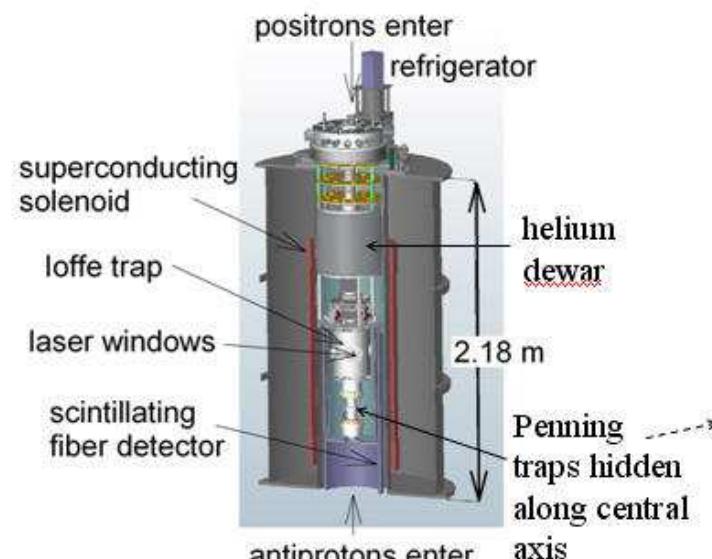
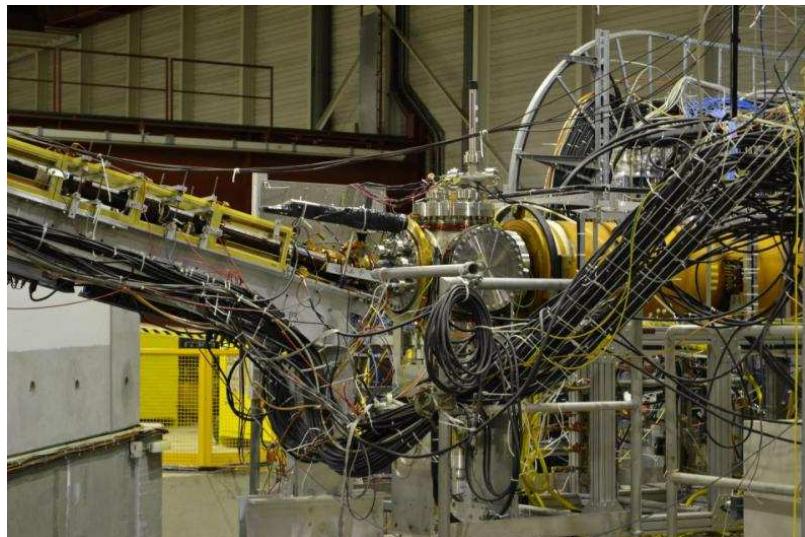
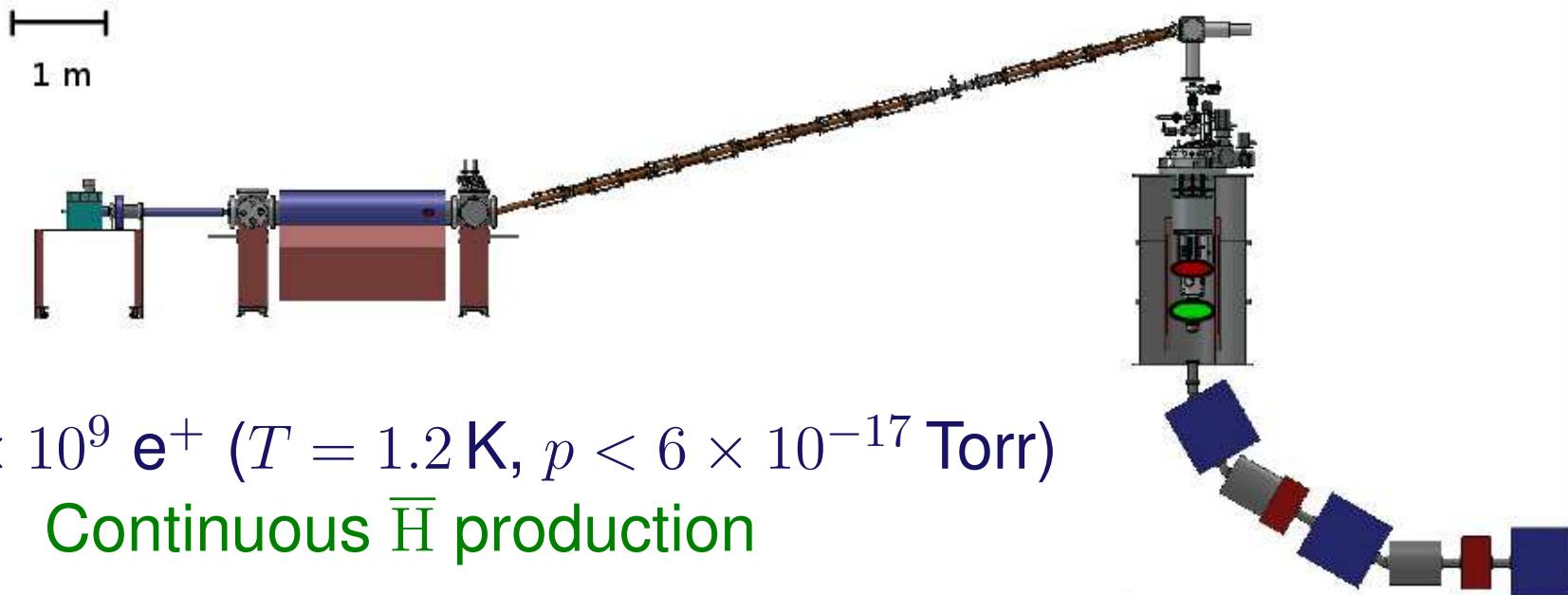
*Characterization of the $1S-2S$ transition in antihydrogen,
Nature* 557 (2018) 74.

Author contributions This experiment was based on data collected using the ALPHA-2 antihydrogen trapping apparatus, designed and constructed by the ALPHA Collaboration using methods developed by the entire collaboration. The entire collaboration participated in the operation of the apparatus and the data-taking activities. The laser and internal cavity system was conceived, implemented, commissioned and operated by W.B., N.M., J.S.H., S.E., C.Ø.R., S.A.J., C.L.C., B.X.R.A. and G.S. F.R., C.Ø.R., J.F. and N.M. developed the simulation program for laser interaction with magnetically trapped atoms. Analysis of the spectral line shapes was done by C.Ø.R., N.M. and J.S.H. Detailed analysis of the antiproton annihilation detector data was done by J.T.K.M. and A.O. Implementation of the microwave system and analysis of the microwave data was done by T.F. and M.E.H. The positron accumulator is the responsibility of C.J.B., M.C., C.A.I. and D.P.v.d.W. The manuscript was written by J.S.H., N.M., C.Ø.R., S.A.J. and J.T.K.M., with help from A.O., C.L.C. and S.E. The manuscript was then edited and improved by the entire collaboration.

Reviewer information *Nature* thanks D. Horvath, K. Jungmann and the other anonymous reviewer(s) for their contribution to the peer review of this work.



ATRAP: Antimatter trap



Antimatter gravity



I read a book on anti-gravity



I couldn't put it down!

Negative mass \Rightarrow repulsive gravity??

95 % of nucleon mass is energy, small grav. diff. between H and \bar{H}

Not *CPT*: weak equivalence principle



Antimatter gravity = supergravity?

Einstein's gravity does not depend on baryon charge

One needs a vector force for that. Supergravity?

Exact supersymmetry cancels even p-p gravity \Rightarrow broken

Trouble: \bar{p} weight on Earth $\simeq e + -e + -$ at 12 cm

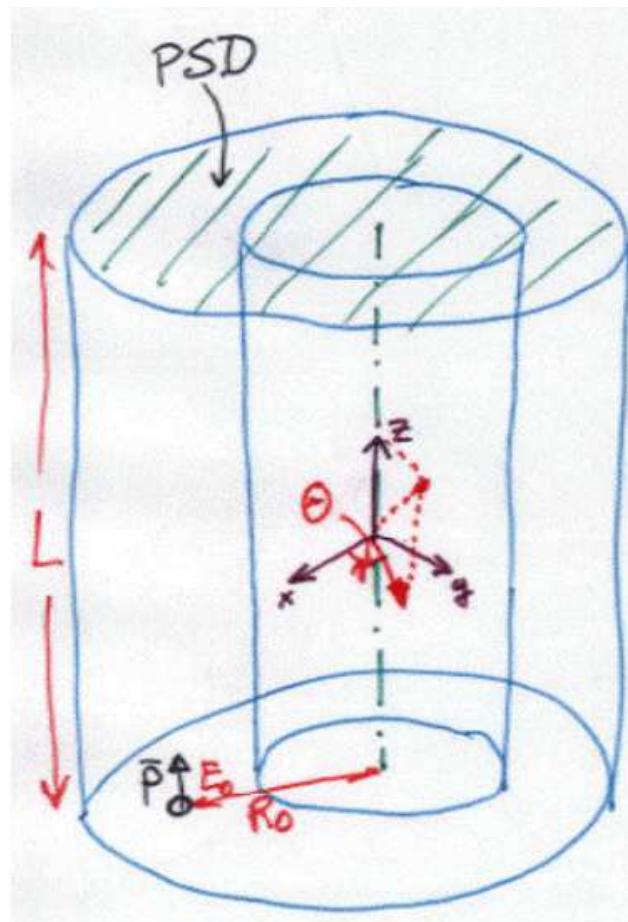
For a 1 % measurement no point charge at 1 m!

Separate gravity effect from the initial motion of \bar{p} :

Magnetron method: J. Eades et al, CERN/PSCC/98-30,
1989



Antiproton gravity: magnetron method?



In cylindr. coord.:

$$B_r = B_z = 0;$$

$$B_\theta = \frac{m}{r}$$

$$v_z = \text{const} = v_z^0$$

Magnetron motion
around z

Cyclotron motion
around B -lines

Radial drift:

$$\vec{v}_d = \frac{m\vec{g} \times \vec{B}}{|g|B^2}; \Delta R \approx R_0 \frac{mg}{E_0}$$

(DH, RIKEN, 1998)

Simulations, Pisa, 1993:

Patch effect kills it.

Random potentials on clean
metal surface:

$\pm 0.01 \dots 0.1 V$ on $1 - 10 \mu m$
patches.

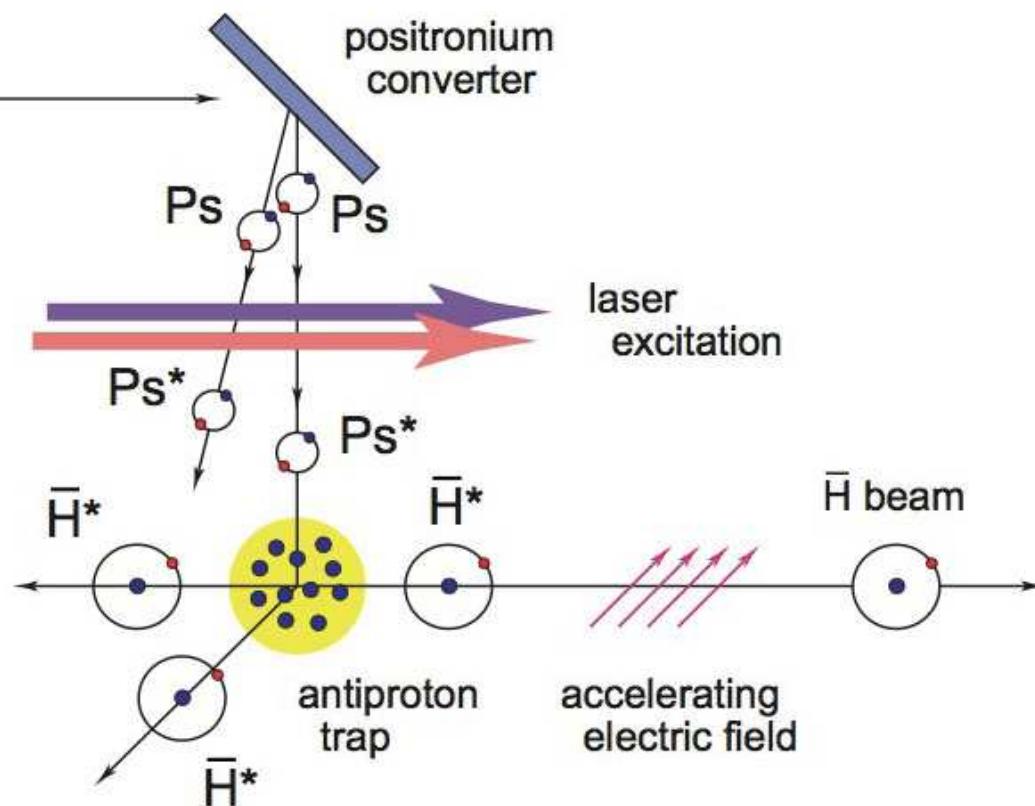
For fast \bar{p} it averages, but no
grav. drift

For slow \bar{p} initial patch
condition dominates.

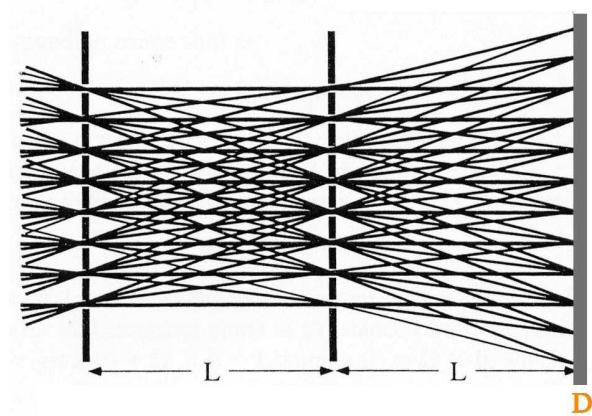
Must use neutral probe:
antihydrogen

AEGIS: antimatter gravity

Antihydrogen Experiment: Gravity, Interferometry, Spectroscopy (in preparation, 77 authors)



Moiré deflectometry:
gravitational falling of
collimated \bar{H}
as compared to light

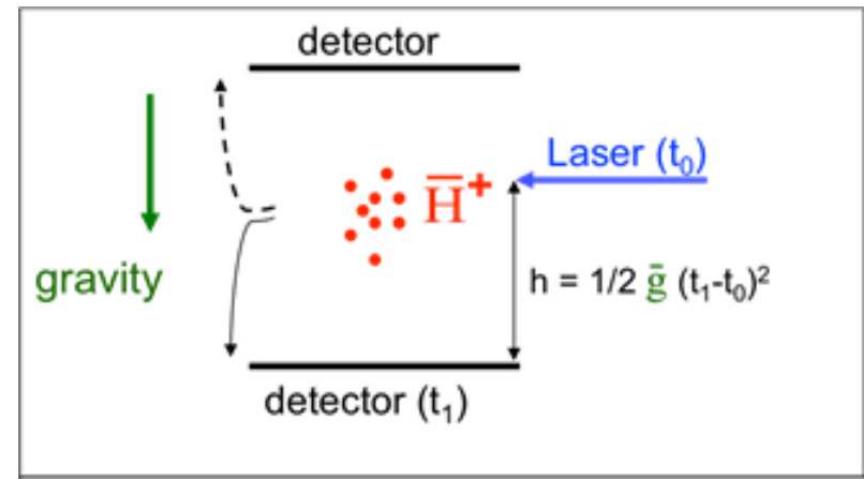
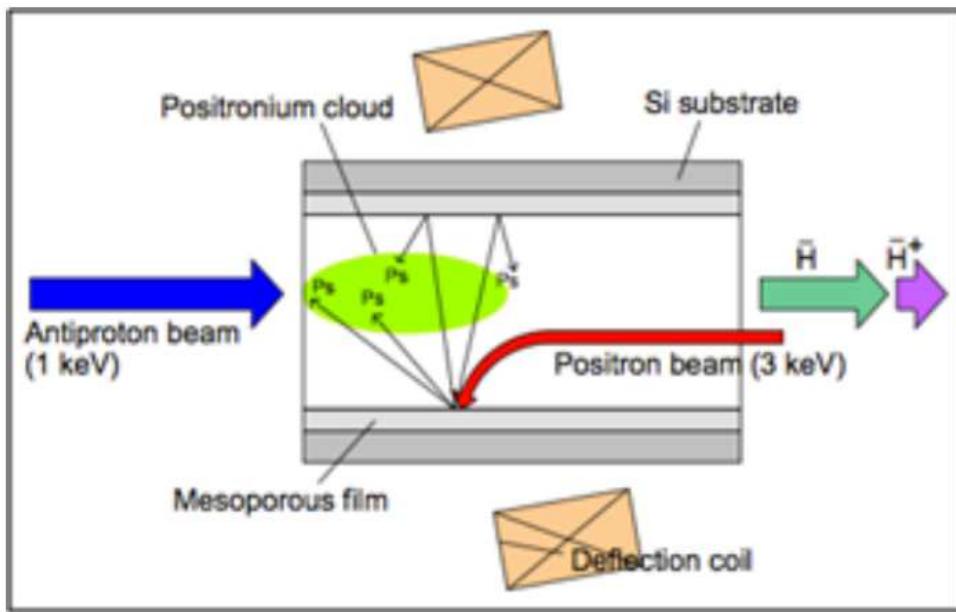


Stark acceleration (electric dipole in inhom. E-field) of
excited \bar{H}



GBAR

Gravitational Behaviour of Antihydrogen at Rest (in preparation)

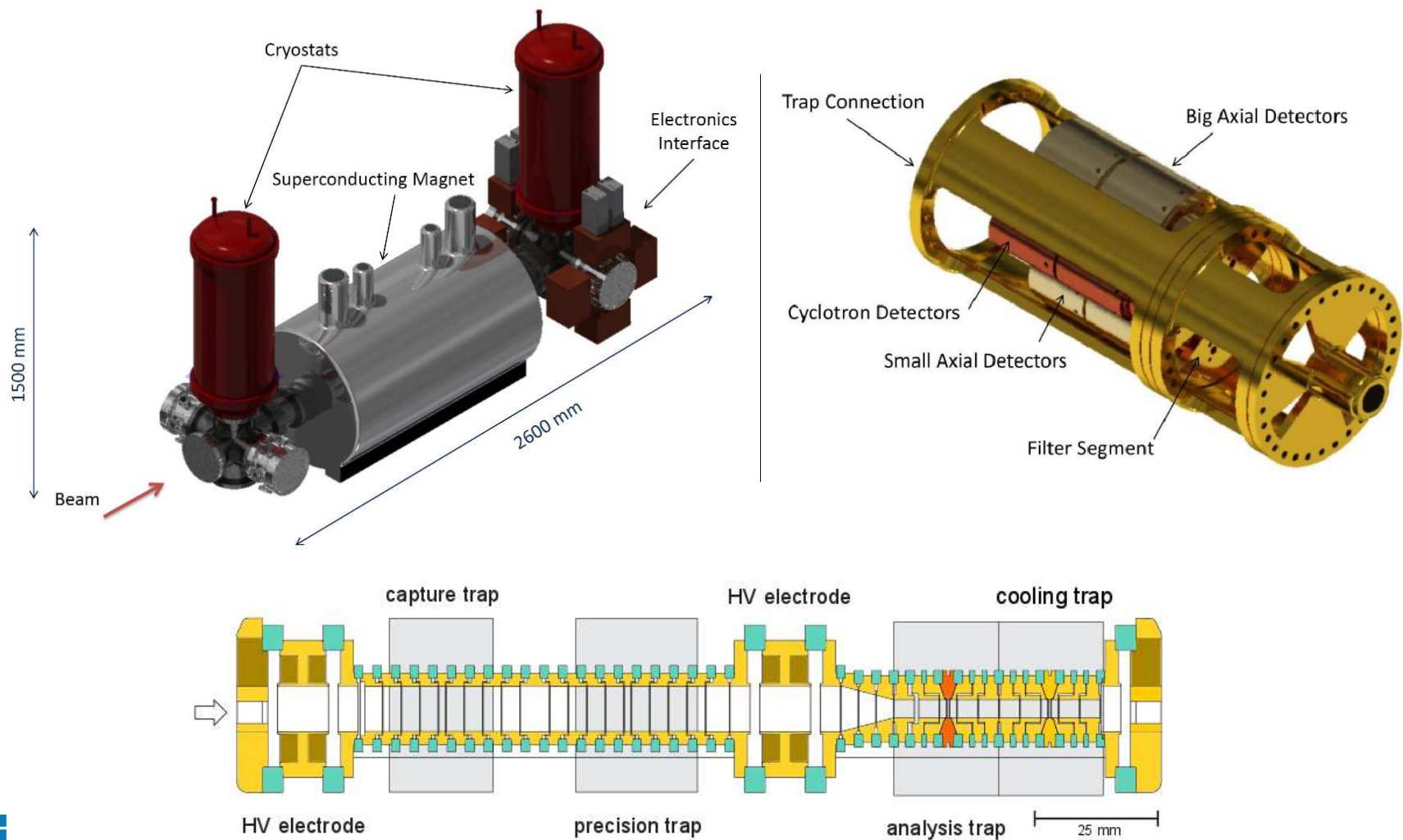


$\bar{p} + Ps \rightarrow \bar{H}; \bar{H} + Ps \rightarrow \bar{H}^+ \text{ (cooling); back to } \bar{H}: \text{ let it fall}$



BASE: Baryon Antibaryon Symmetry Experiment

Direct high-precision measurement of the magnetic moment of a single antiproton stored in a cryogenic Penning trap



(in preparation)

Antihydrogen beam

ASACUSA: MUSASHI



Monoenergetic
Ultra
Slow
Antiproton
Source for
High-precision
Investigations

Musashi Miyamoto self-portrait ~ 1640

5.8 MeV \bar{p} injected into RFQ

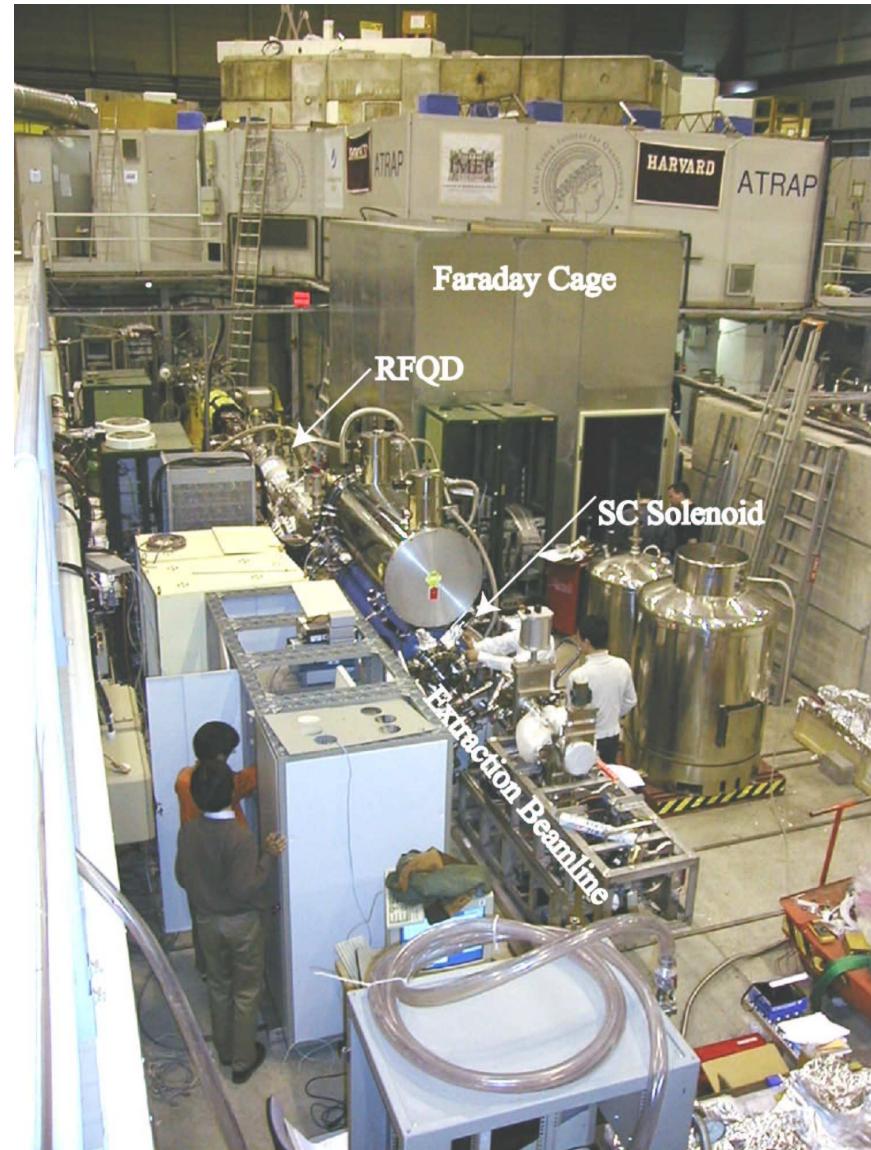
100 keV \bar{p} injected into trap

$10^6 \bar{p}$ trapped and cooled (2002)

~ 350000 slow \bar{p} extracted (2004)

Cold \bar{p} compressed in trap (2008)

$(5 \times 10^5 \bar{p}, E = 0.3 \text{ eV}, R = 0.25 \text{ mm})$

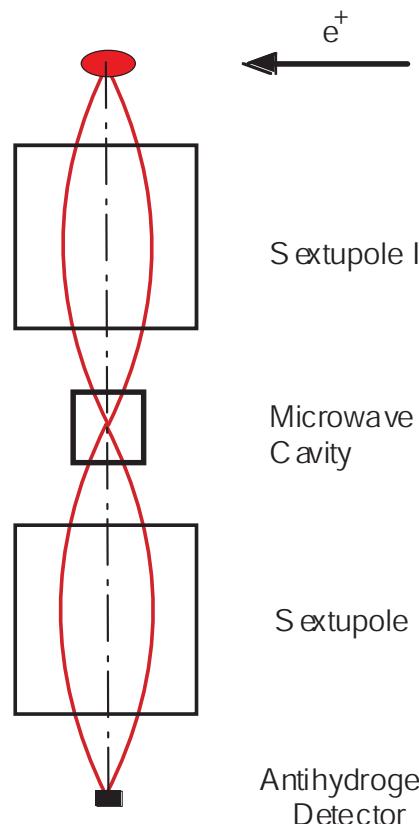


\bar{H} -beam: N. Kuroda *et al.*, Nature Commun. 5 (2014) 3089.

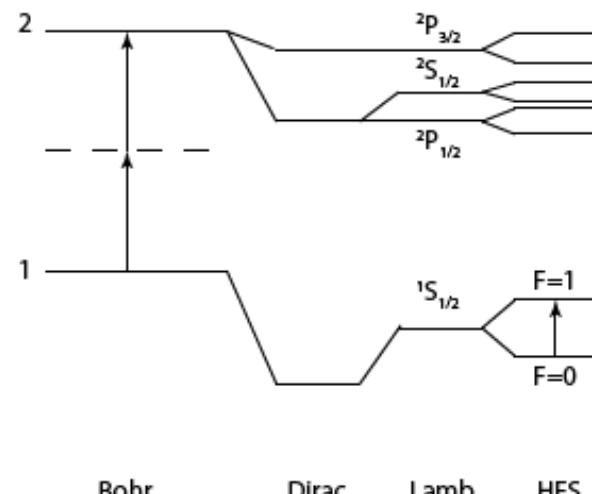


Spectroscopy with \bar{H} -beam

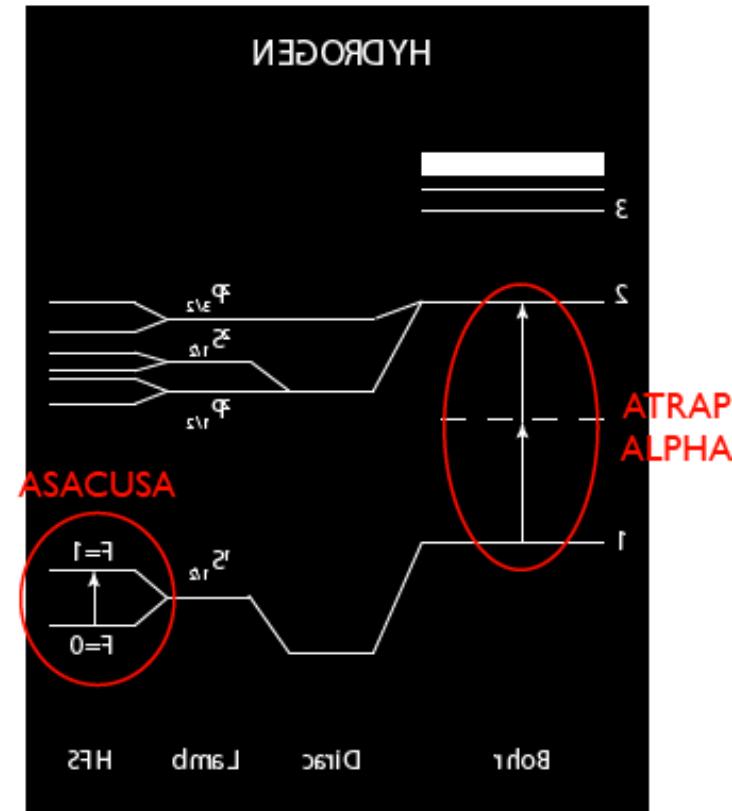
antiproton and positron
Trap / Recombination



HYDROGEN



HYDROGEN



\bar{H} -beam path: polariser, resonator, analyser
Analogy to polarised light

R.S. Hayano, M. Hori, D. Horváth, E. Widmann, Rep. Progr. Phys. 70 (2007) 1995.

Extra Low ENergy Antiprotons (ELENA)

Physics Motivation

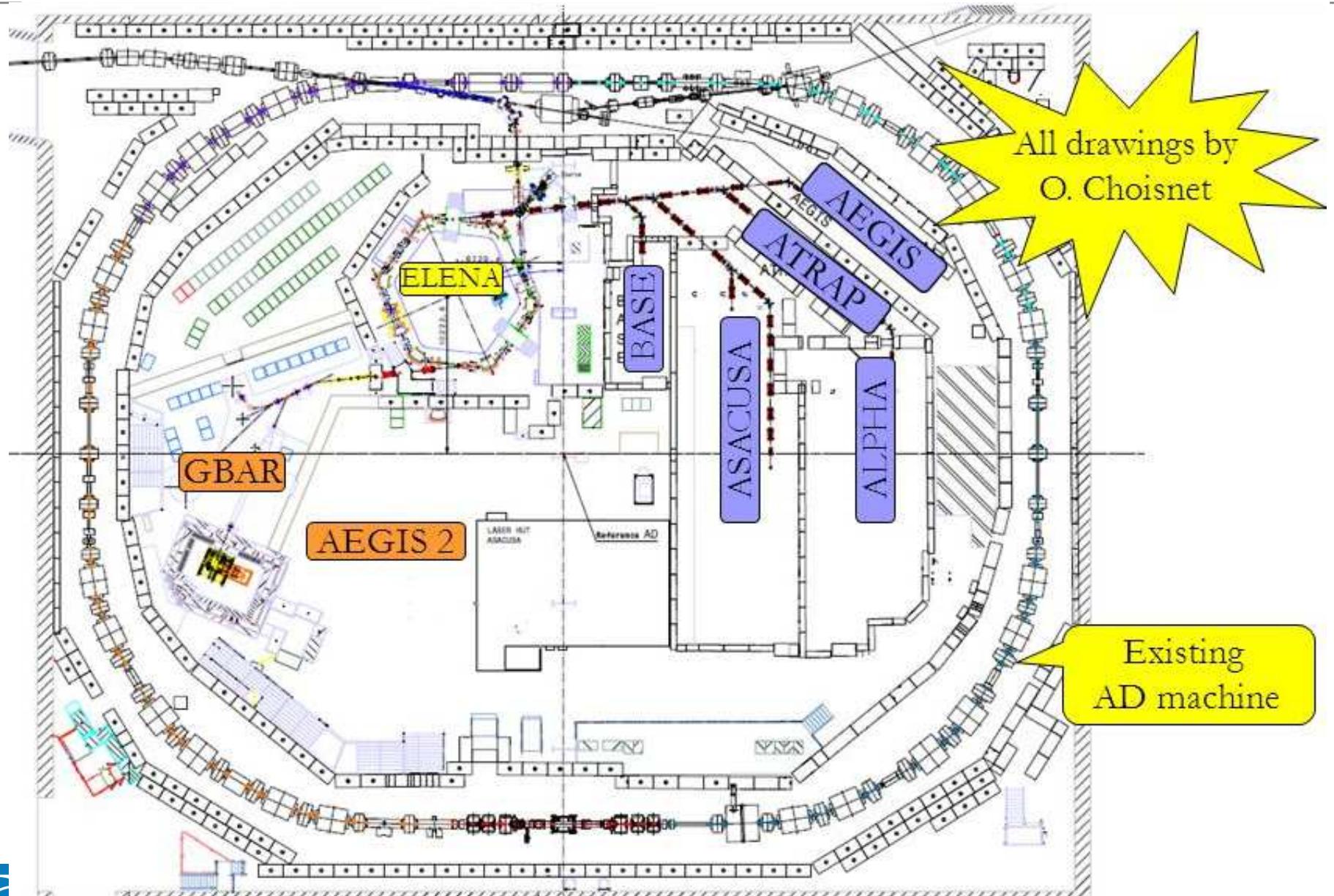
- Test the Standard Model and General Relativity for antimatter
- Test SM extensions for antimatter (Lorentz-violation, black holes, new interactions, ...)
- Stringent CPT tests with antihydrogen
- Antimatter gravity measurement (weak equivalence)
- Added precision for physical constants (CODATA) assuming CPT invariance

All existing AD experiments profit, new ones made possible (gravity, X-rays, nuclear studies)

Dániel BARNA (Wigner RCP) participated in its development



ELENA at the AD: plan



F. Butin / ELENA collaboration, 2017.

Antimatter in Space

AMS-2: Alpha Magnetic Spectrometer
to discover antimatter (anti-helium!) and
dark matter

Mass: 8500 kg,
1200 kg perm. magnet

Father: Sam Ting, cost: 2 G\$

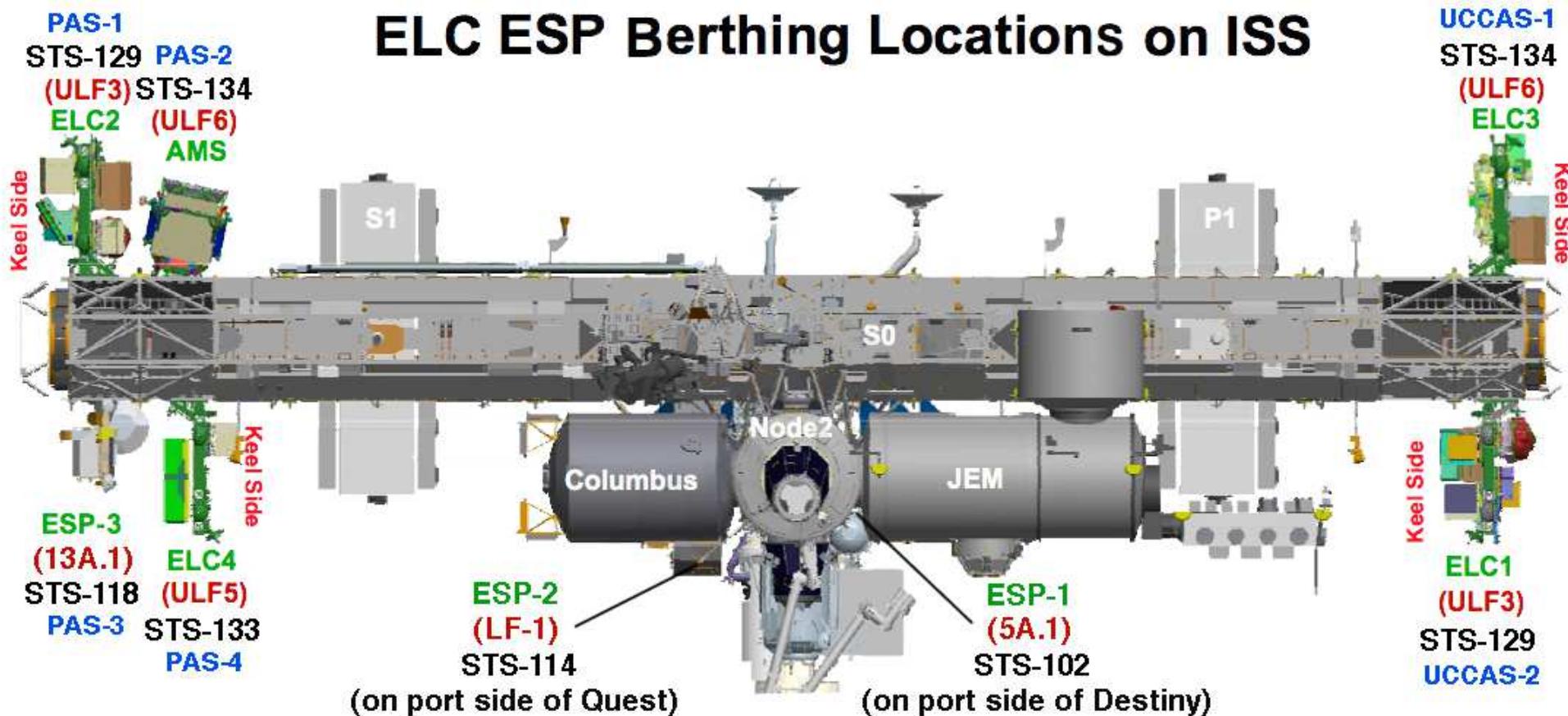
Construction: CERN

Launch: May 2011, USA

Control room: CERN



AMS-2: Alpha Magnetic Spectrometer



First results (2013-14):

No antihelium observed.

High energy positrons everywhere.

Could come from dark matter or pulsars.

AMS2 will collect data for 10–15 years.

ASACUSA: Mass of the antiproton

Proton's well (?) known:

$$m(p)/m(e) = 1836.15267245(75)$$

$$q(e) = 1.602176565(35) \times 10^{-19} \text{ C}$$

Precision: $4 \cdot 10^{-10}$ and $2 \cdot 10^{-8}$

Relative measurements: proton vs. antiproton

Cyclotron frequency in trap $\rightarrow q/m$

TRAP \Rightarrow ATRAP collaboration

Harvard, Bonn, München, Seoul

\bar{p} and H^- together $\Rightarrow 10^{-10}$ precision

Atomic transitions:

$$E_n \approx -m_{\text{red}} c^2 (Z\alpha)^2 / (2n) \rightarrow m \cdot q^2$$

PS-205 \Rightarrow ASACUSA collaboration

Tokyo, Brescia, Budapest, Debrecen, Munich, Vienna

Atomic
Spectroscopy
And
Collisions
Using
Slow
Antiprotons

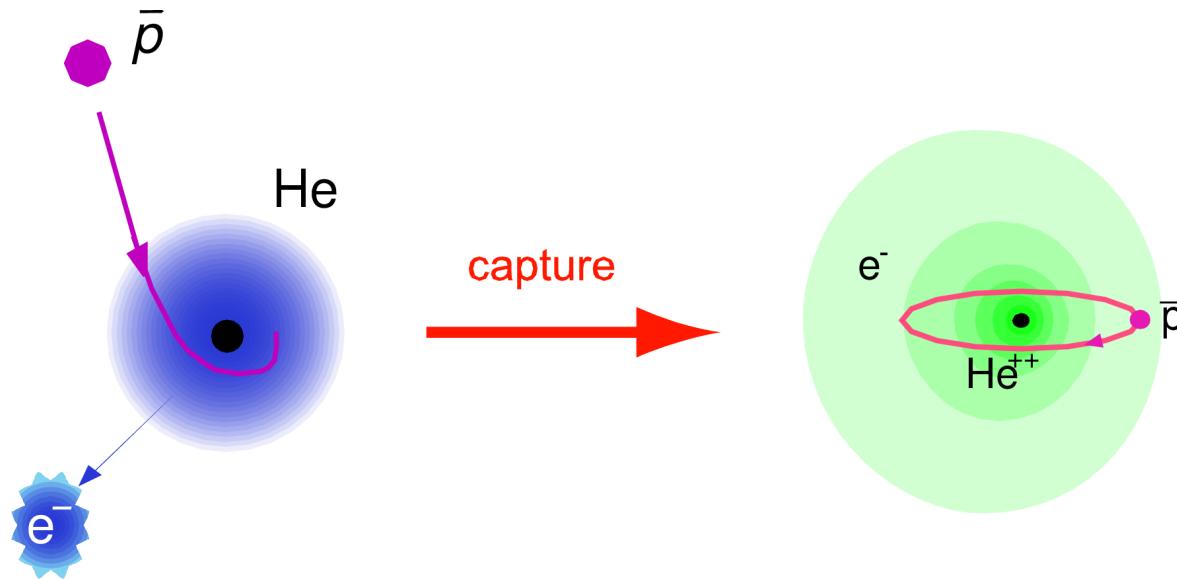


Asakusa, Tokyo



Metastable hadronic atoms

In matter (gas, liquid, solid) $\tau(\text{hadron}) \sim 1 \text{ ps}$
except $\sim 3\%$ of $X^- \text{He}$: K^-, π^- : decay lifetime; \bar{p} : $3-4 \mu\text{s}$



Metastable 3-body system
Auger suppressed, slow radiative transitions only

Electron *cloud* protects \bar{p} against collisions

Electron tightly bound: $1S$

$\bar{p}\text{He}$: $n \sim 40$, $l \sim n - 1$, Rydberg state

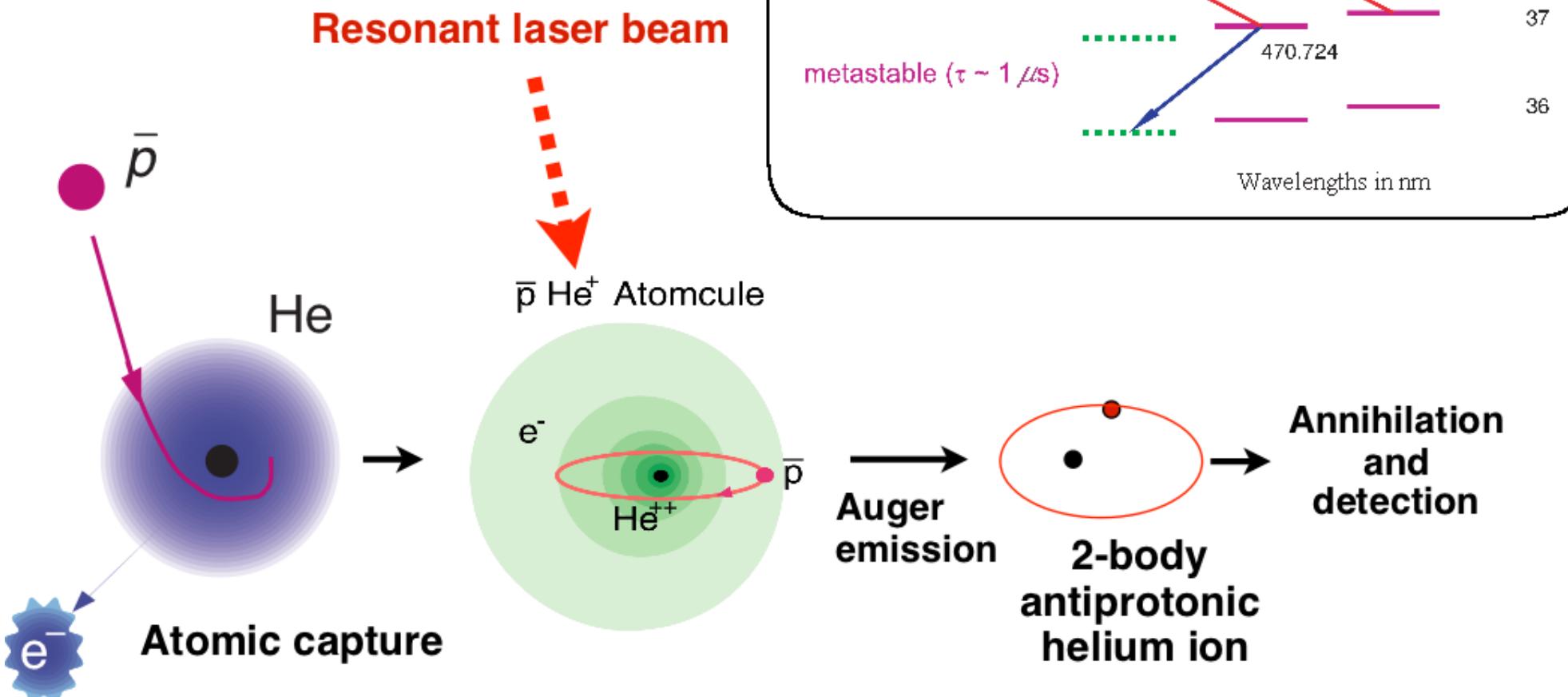


\bar{p} -He⁺: spectroscopy motivation

- Vladimir Korobov calculates \bar{p} transition frequencies in \bar{p} -He⁺ with the precision of $\sim 10^{-9}$
- Determination of antiproton-to-electron mass ratio to 1.3×10^{-9} .
→ Dimensionless fundamental constant of nature.
- Determination of electron mass in a.u. to 1.3×10^{-9}
→ One of the data points for CODATA2010 average.
- When combined with cyclotron frequency of antiprotons in a Penning trap measured by the TRAP collaboration, comparison of antiproton and proton mass and charge to 7×10^{-10}
→ Particle Data Group: CPT consistency test.



Laser spectroscopy of antiprotonic helium



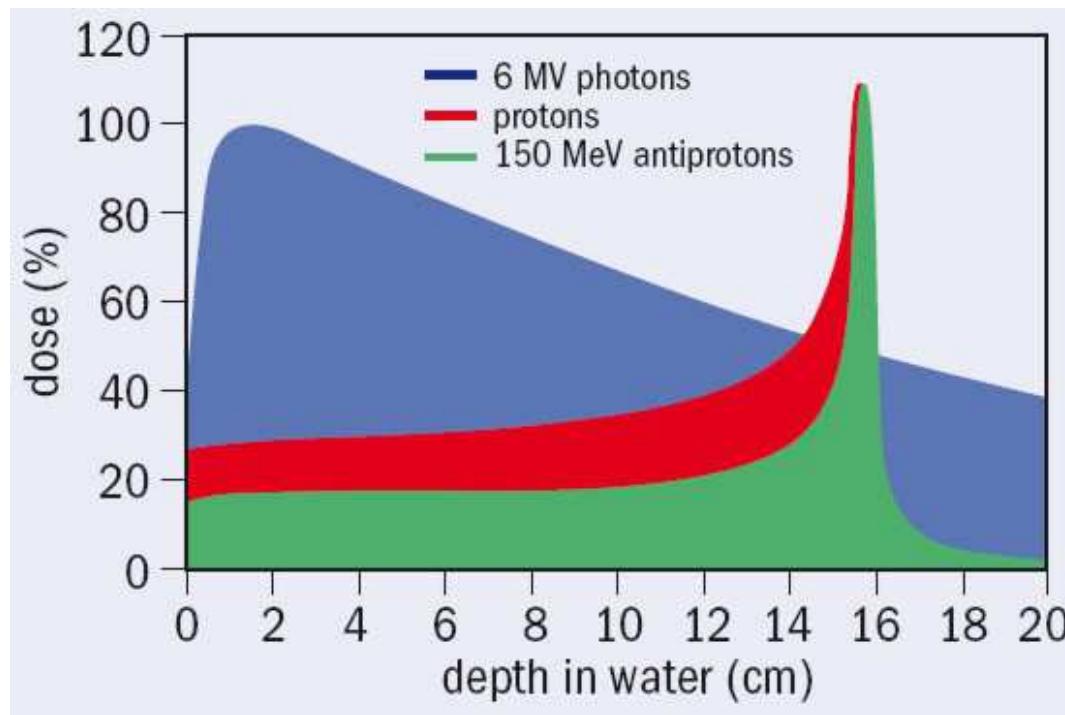
Induce transition between long-lived and short-lived states



Force prompt annihilation

ACE: Antimatter Cell Experiment

Cancer therapy research (USA) at AD of CERN



Advantage: Antiprotons lose energy in very small volume, choosing the right energy concentrates damage in tumor.

Disadvantage: Antiprotons are very expensive and annihilation radiation damages as well.



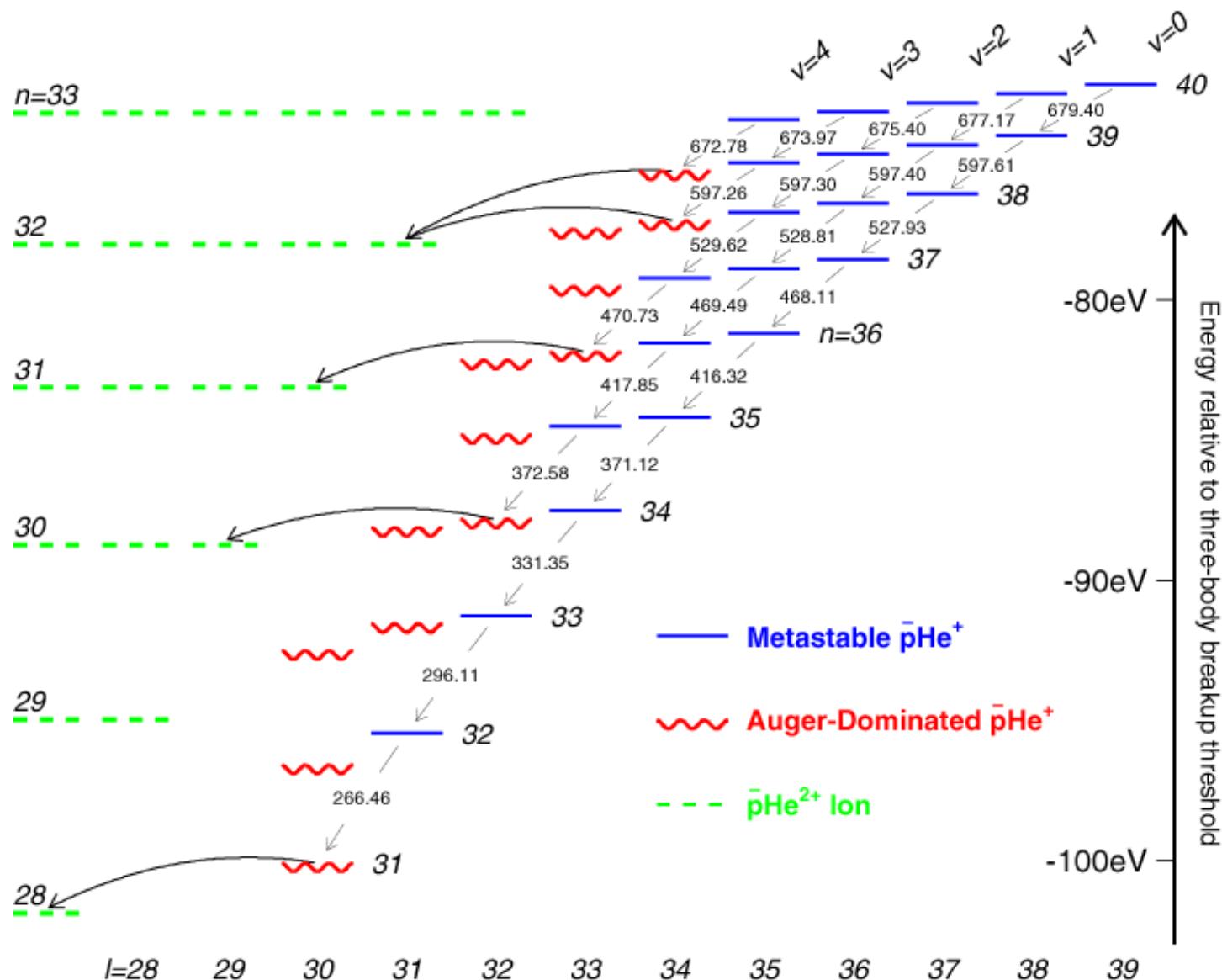
Thanks for your attention



Spare slides for discussion



Energy levels of $\bar{p}\text{He}^4$



Level energies in eV, transition wavelengths in nm

MUSASHI: slow antiproton beam



Monoenergetic
Ultra
Slow
Antiproton
Source for
High-precision
Investigations

Musashi Miyamoto self-portrait \sim 1640

5.8 MeV \bar{p} injected into RFQ

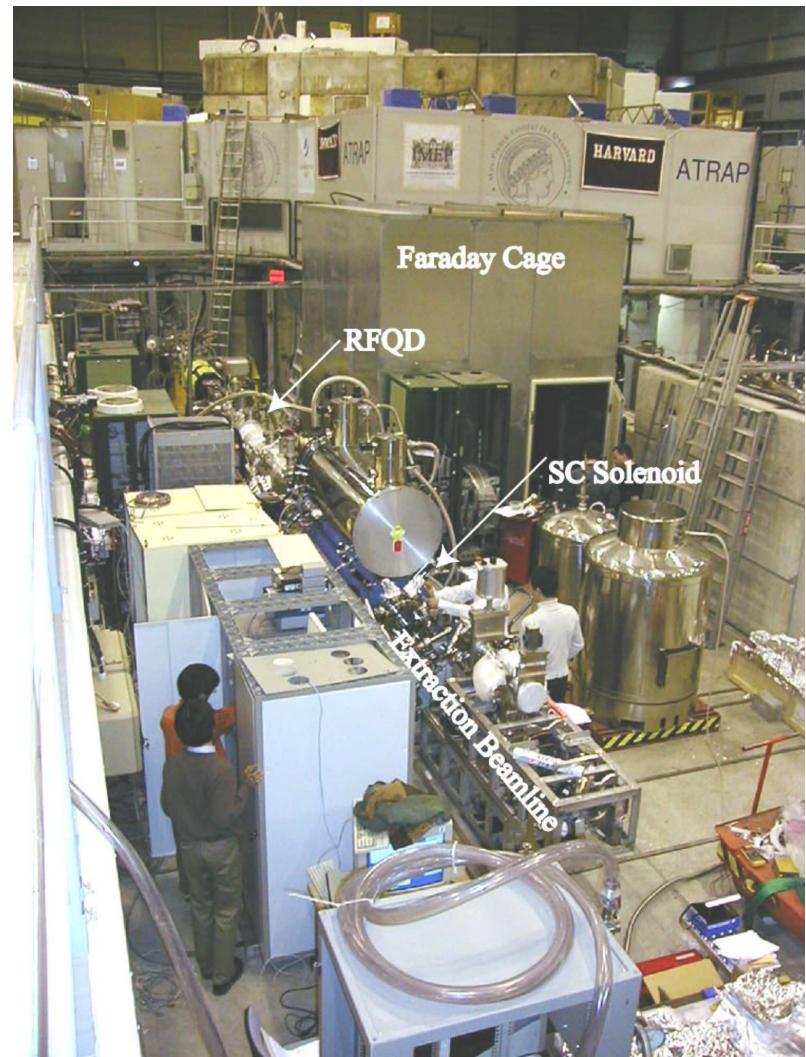
100 keV \bar{p} injected into trap

10^6 \bar{p} trapped and cooled (2002)

~ 350000 slow \bar{p} extracted (2004)

Cold \bar{p} compressed in trap (2008)

$(5 \times 10^5 \bar{p}, E = 0.3 \text{ eV}, R = 0.25 \text{ mm})$



N. Kuroda,...D. Barna, D. Horváth, Y. Yamazaki: Phys. Rev. Lett. 100 (2008) 203402.



Two-photon spectroscopy

In low density gas main precision limitation:
thermal Doppler broadening even at $T < 10$ K

Excite $\Delta\ell = 2$ transition with 2 photons

Two counterpropagating photons with $\nu_1 \sim \nu_2$
eliminate 1st order Doppler effect

Laser linewidth should not overlap with resonance

M. Hori, et al., A. Sótér, D. Barna, A. Dax, R.S. Hayano, S. Friedreich, B. Juhász, T. Pask, E. Widmann, D. Horváth, L. Venturelli, N. Zurlo:

Two-photon laser spectroscopy of pbar-He⁺ and the antiproton-to-electron mass ratio, Nature 475 (2011) 484-488,
Few Body Syst. 54 (2013) 917-922.



Two-photon spectroscopy: parameters

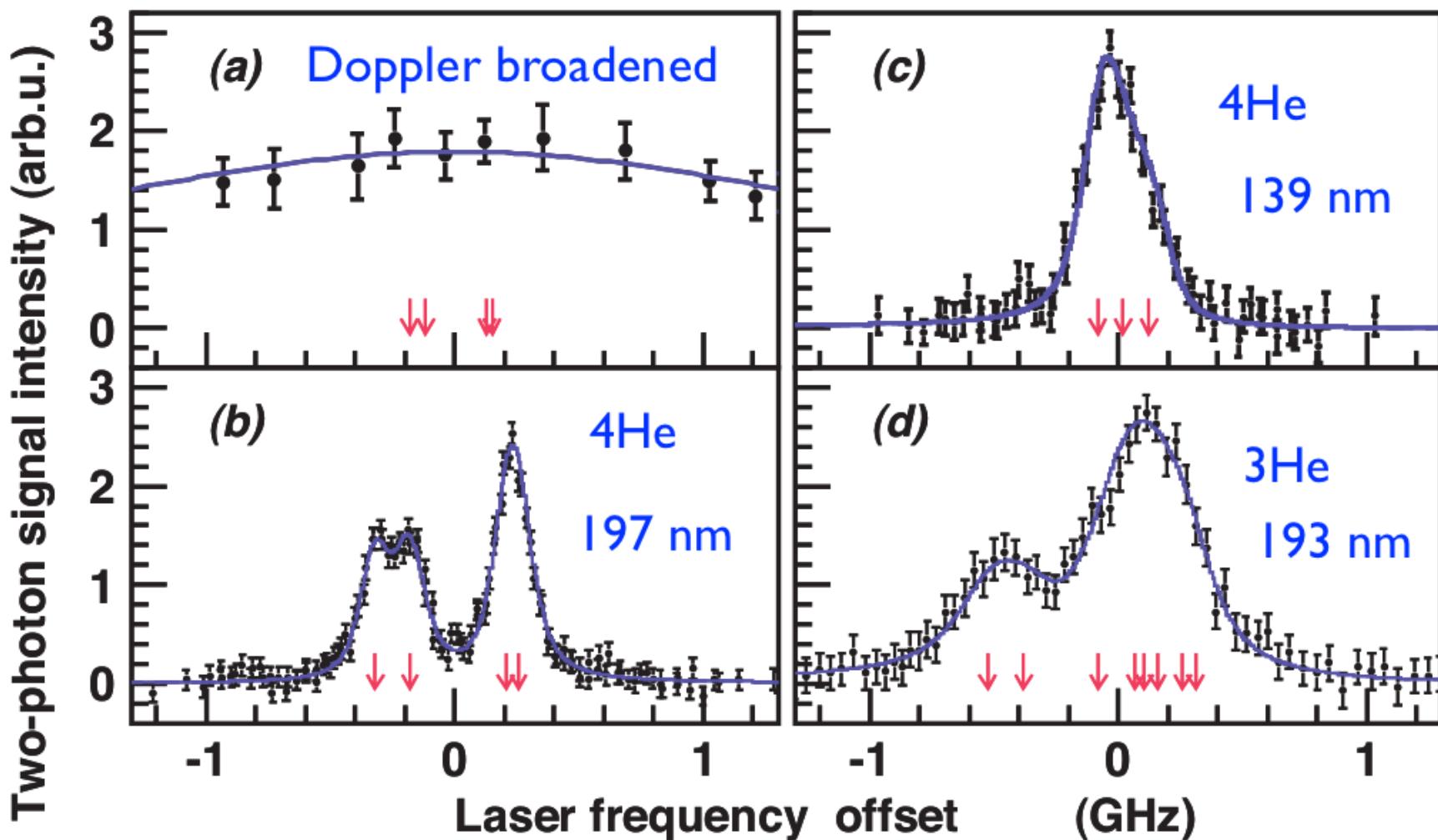
- Precision of lasers: $< 1.4 \times 10^{-9}$.
- $7 \times 10^6 \text{ p}/\text{pulse}$, $E \approx 70 \text{ keV}$, 200 ns long, Ø20 mm.
- Target: He gas, $T \approx 15 \text{ K}$, $p = 0.8 - 3 \text{ mbar}$
- Laser beams: $\lambda_1 = 417 \text{ nm}$, $\lambda_2 = 372 \text{ nm}$, $P \approx 1 \text{ mJ/cm}^2$
- Transition: (n=36, l=34) \rightarrow (n=34, l=32); $\Delta\nu = 6 \text{ GHz}$
- Measured linewidth: $\approx 200 \text{ MHz}$
- Width: Residual Doppler broadening, hyperfine structure, Auger lifetime, power broadening.

M. Hori, A. Sótér, D. Barna, A. Dax, R.S. Hayano, S. Friedreich, B. Juhász,
T. Pask, E. Widmann, D. Horváth, L. Venturelli, N. Zurlo: „Two-photon
laser spectroscopy of pbar-He⁺ and the antiproton-to-electron mass ratio”

Nature 475 (2011) 484-488



Two-photon spectroscopy: spectra

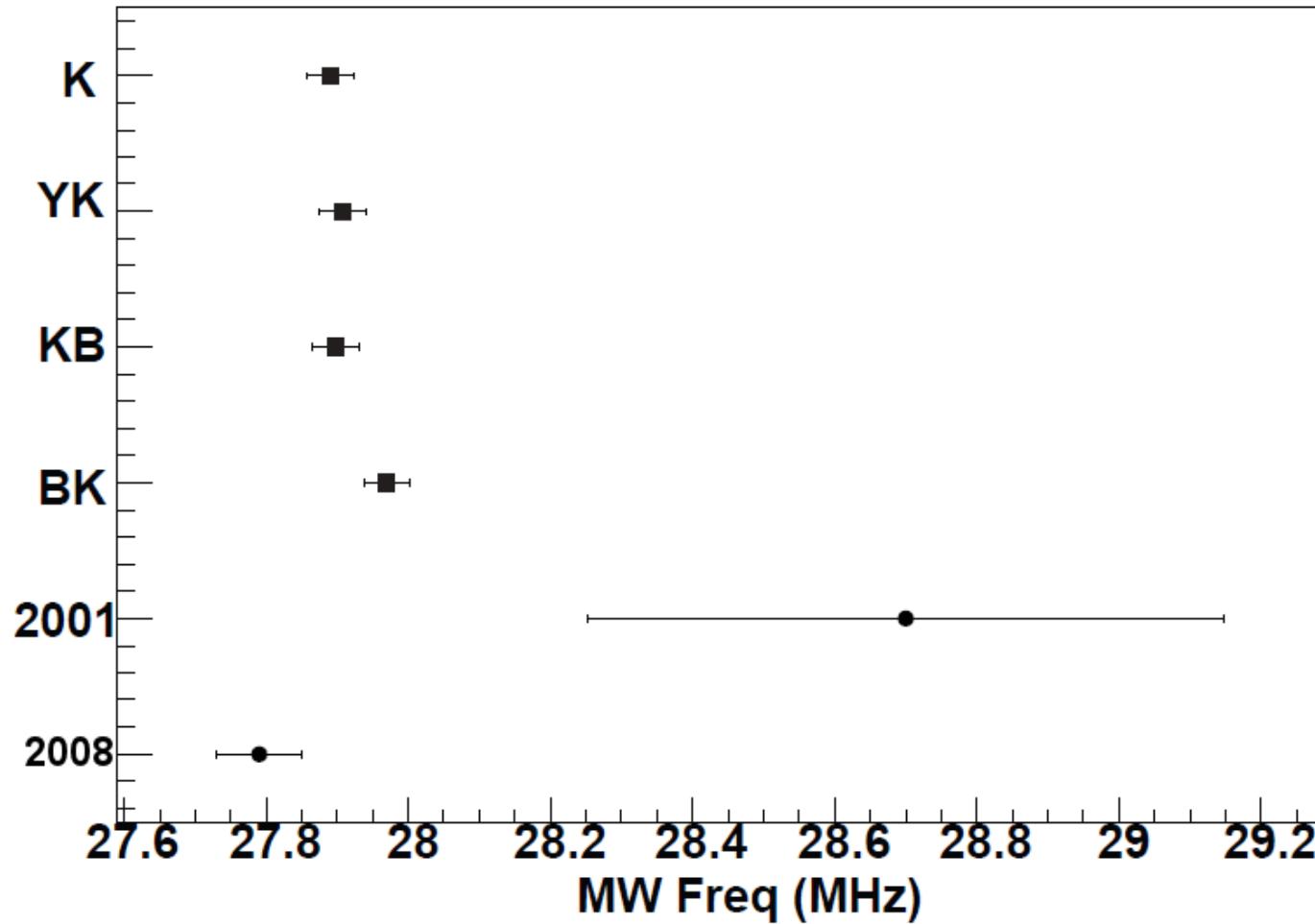


M. Hori et al., Nature 475 (2011) 484-488

Arrows: hyperfine transitions



$\bar{p}^4\text{He}$ HF structure: expt vs. theory

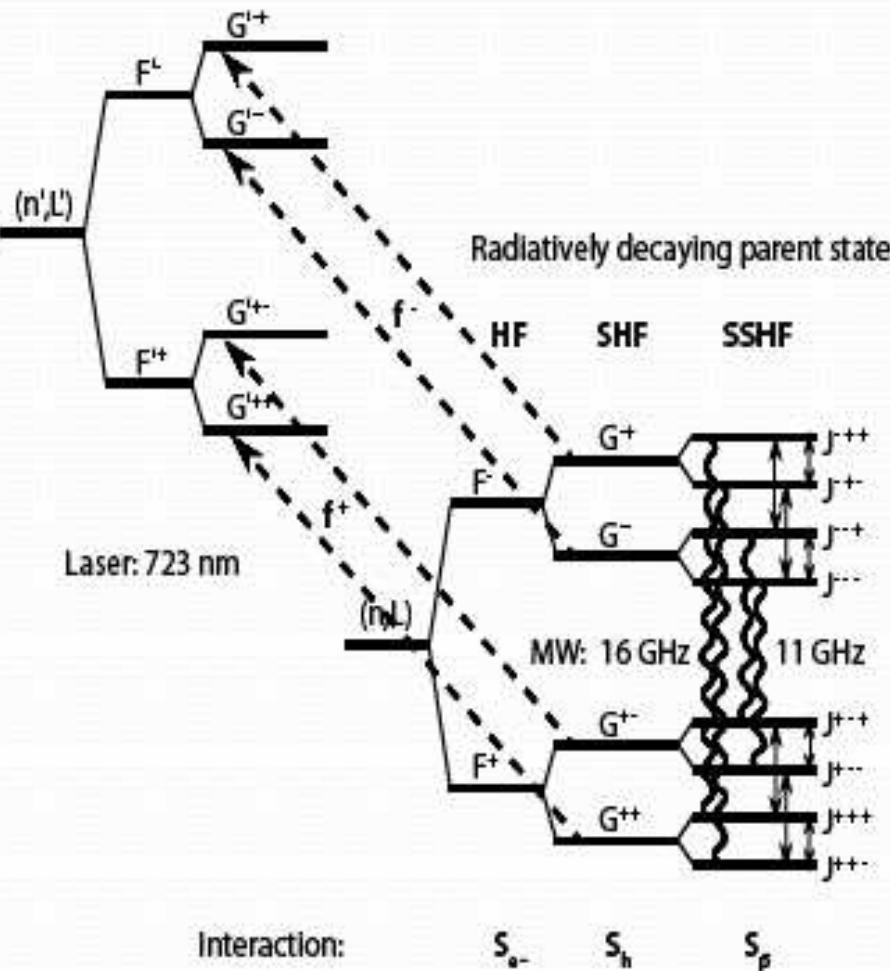


Th. Pask et al., Phys. Lett. B 678 (2009) 55.

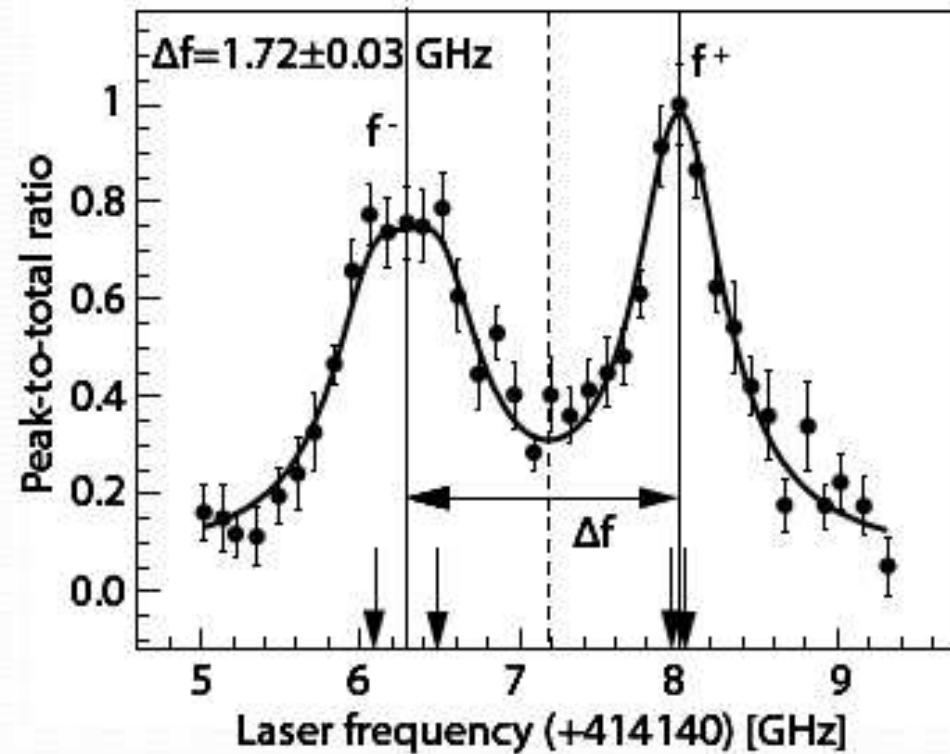


$\overline{p}^3\text{He}$ HF structure: laser scan

Auger decaying daughter state



- verify splitting of laser transition lines
- determine laser resonance frequency



- fit with 4 Voigt functions plus constant for signal background

S. Friedrich et al., Physics Letters B 700 (2011) 1.